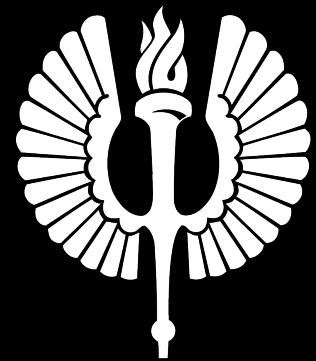


# Core-collapse supernovae in dusty environments



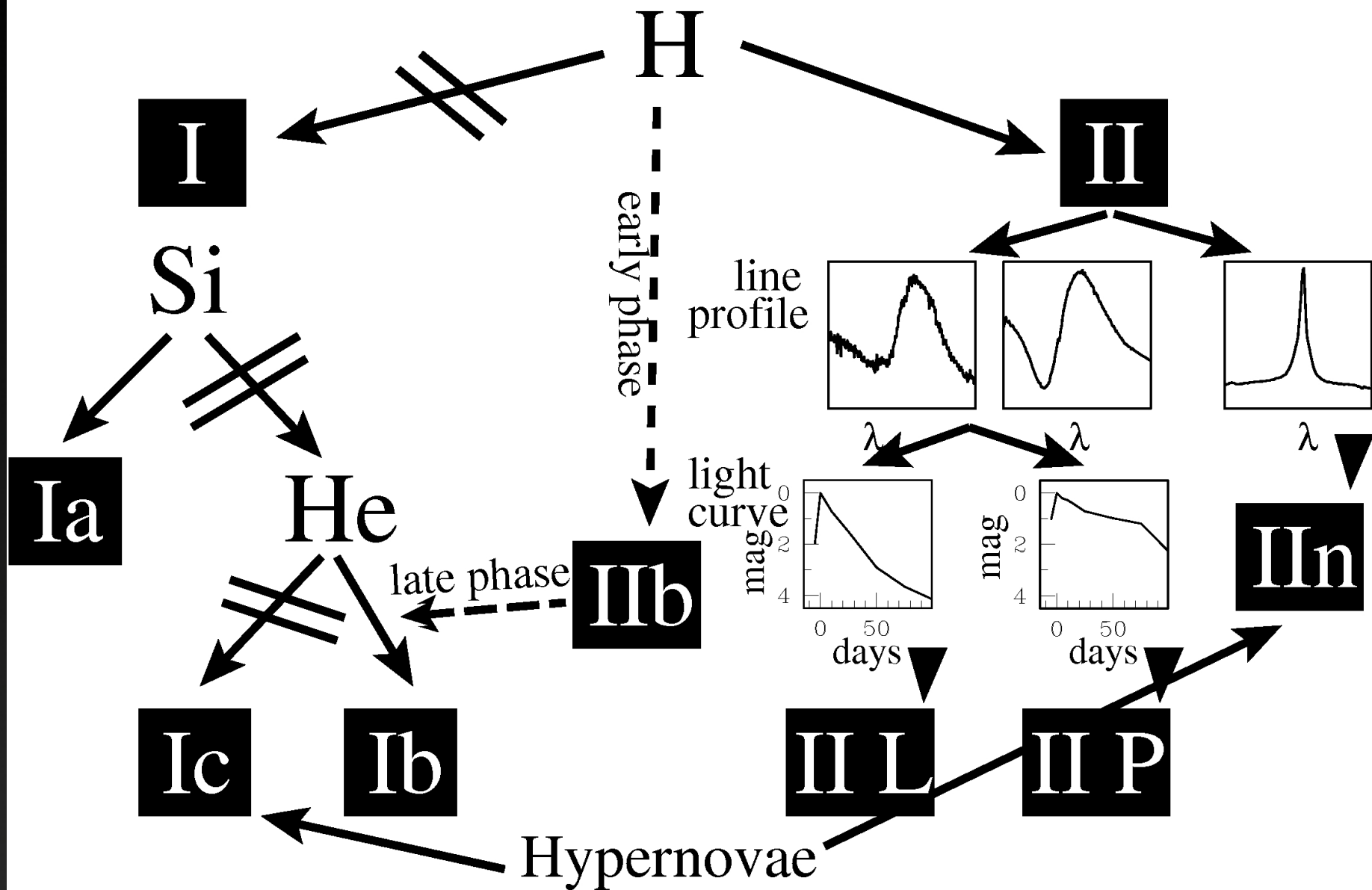
Erkki Kankare

Seppo Mattila, Cristina Romero-Cañizales, Jari Kotilainen,  
Ari Takalo (Turku); Stuart Ryder (AAO); Petri Väisänen (SAAO);  
Miguel Ángel Pérez-Torres, Antxon Alberdi (IAA-CSIC);  
Jens Melinder (Stockholm) et al.



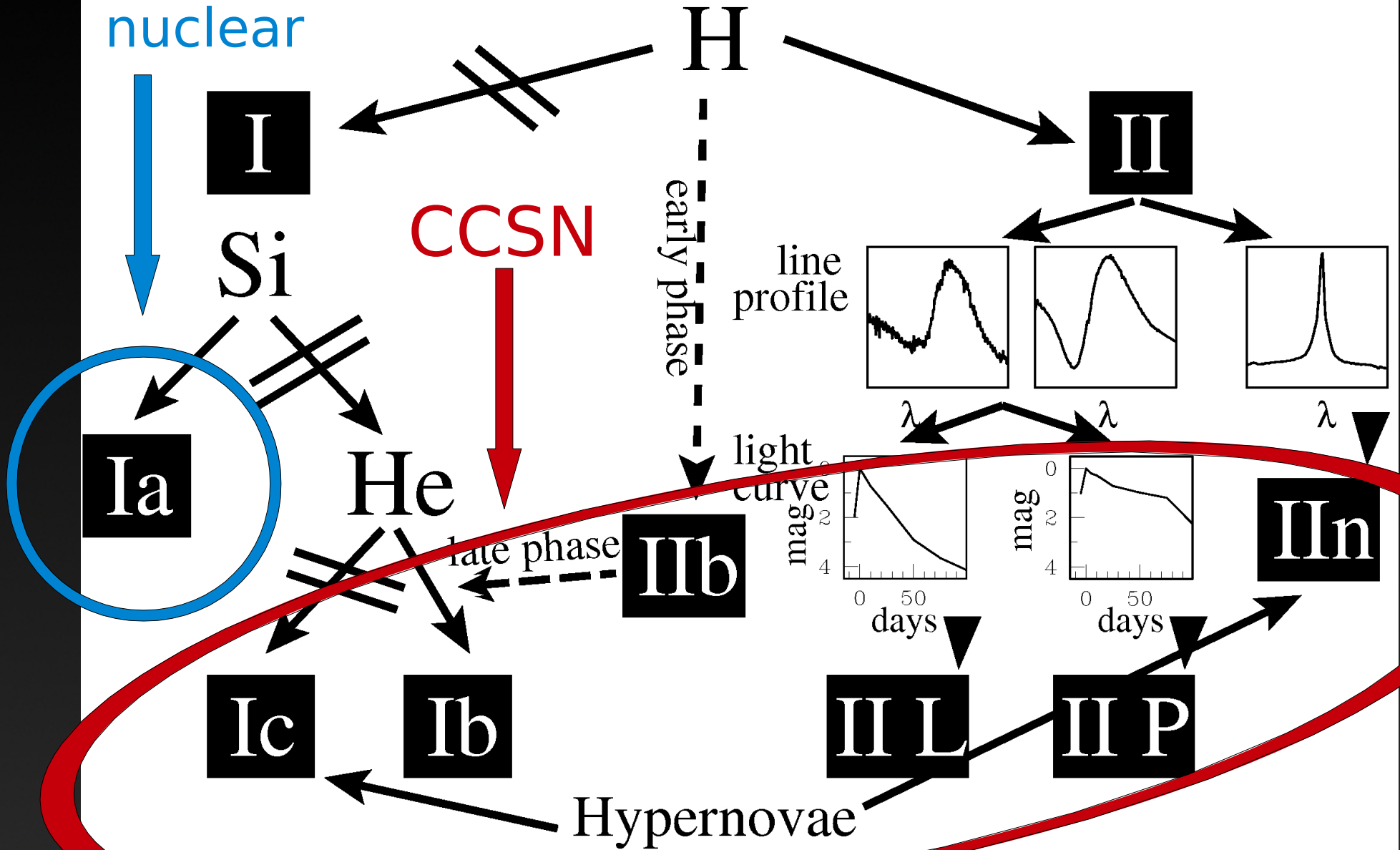
June 4<sup>th</sup> - Astronomers' Days 2012 - Haikko, Porvoo

# Supernova taxonomy

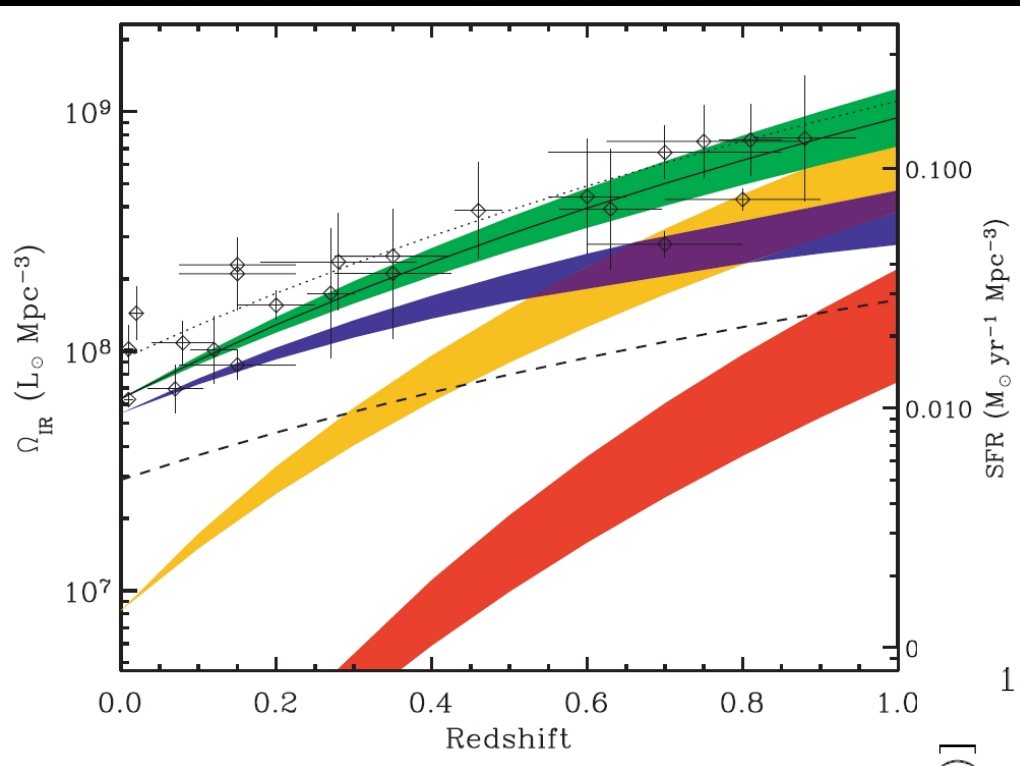


# Supernova taxonomy

Thermo-nuclear



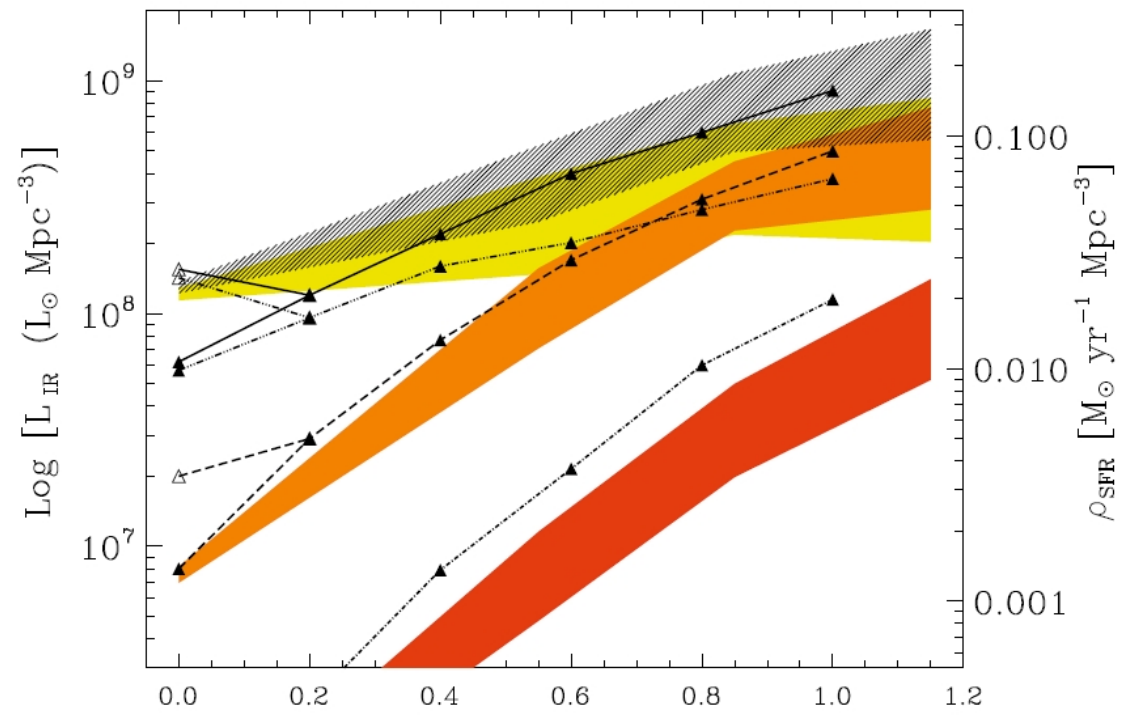
# Supernovae in U/LIRGs



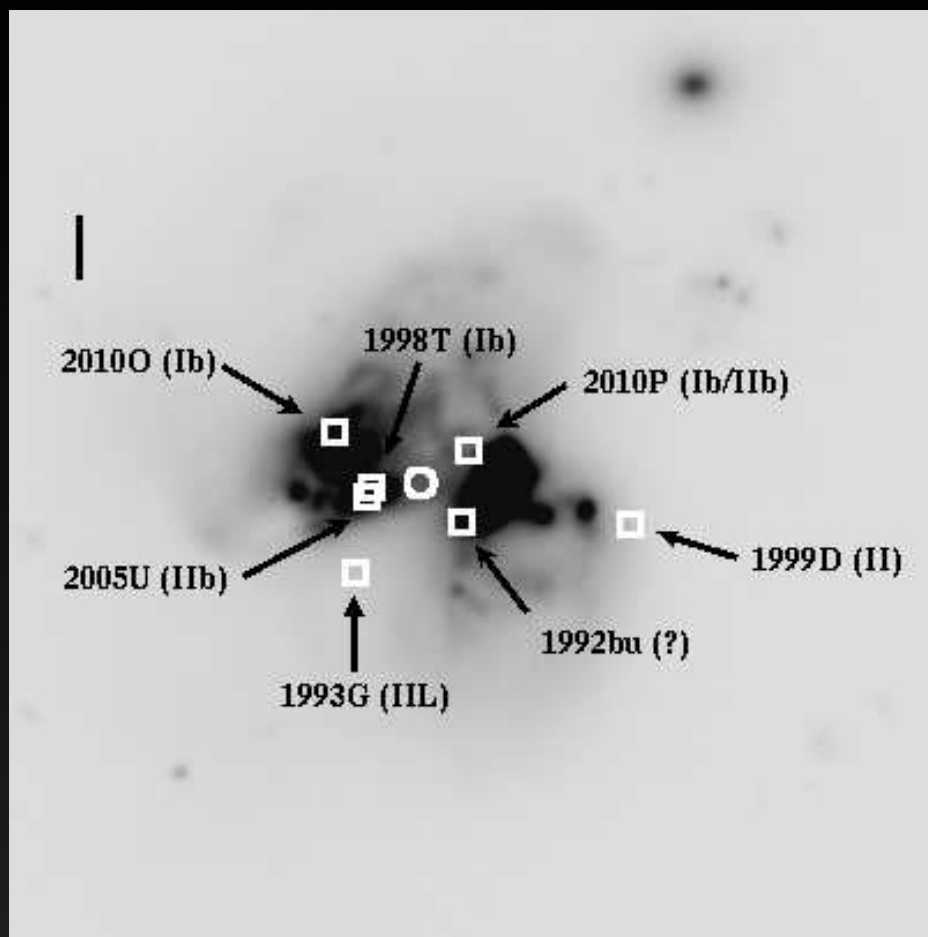
- Luminous ( $L_{\text{IR}} > 10^{11} L_{\odot}$ ) and ultraluminous ( $L_{\text{IR}} > 10^{12} L_{\odot}$ ) infrared galaxies (LIRGs and ULIRGs)
- At redshift  $z \sim 0.7$  U/LIRGs become the dominant source of star formation
- SN rate provides an independent tool to measure star formation

Le Floc'h et al. (2005)

- Dusty, highly obscured galaxies
- SNe can be used to map the extinction distributions in their host galaxies



Magnelli et al. (2009)

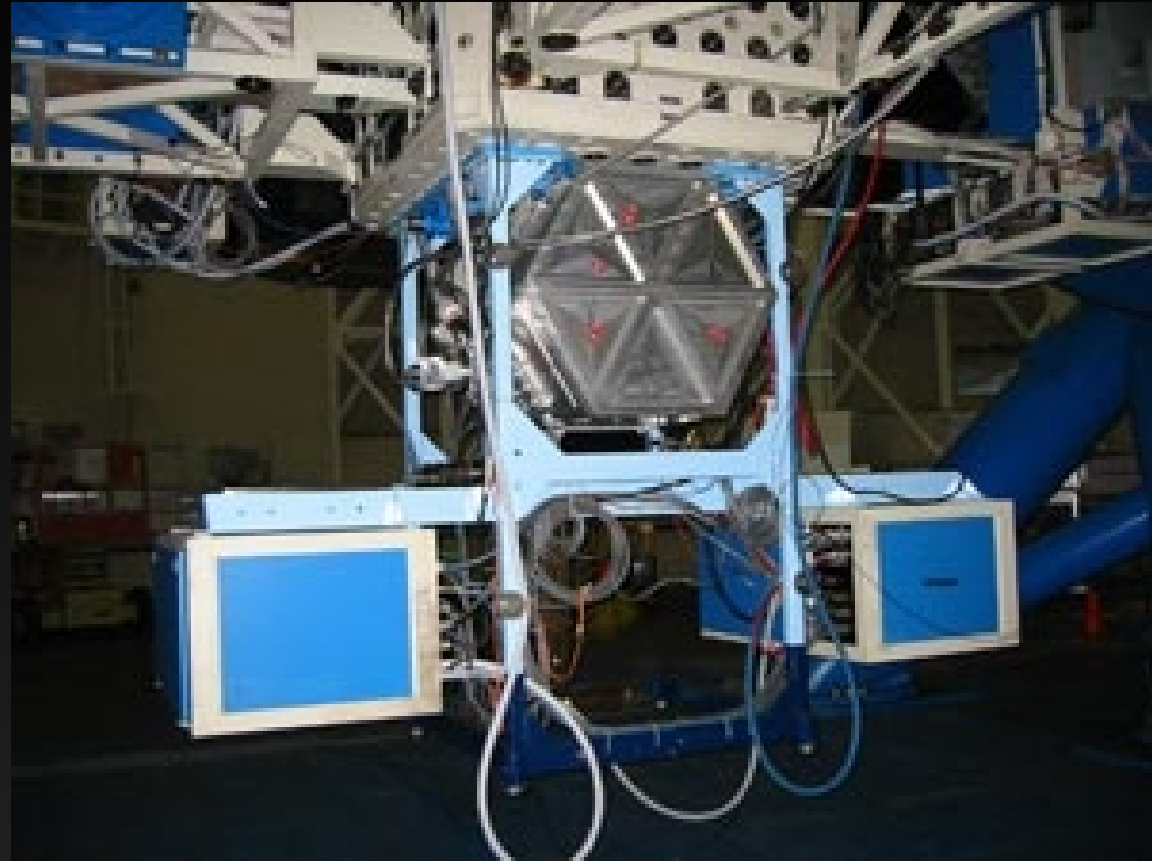


Anderson et al. (2011)

- Potentially different SN population in LIRGs compared to normal spiral galaxies (Anderson et al. 2011)
  - Case study on Arp 299
  - Ratio of stripped envelope (Ib/c, IIb) SNe to normal Type II (IIP/L) SNe higher compared to normal spiral galaxies in the local Universe
- Small number statistics: so far only a low number of SNe have been discovered in U/LIRGs in optical/near-IR
- Dusty, highly obscured environments - especially in the nuclear regions
- Near-IR observations and high spatial resolution required
  - Adaptive optics
  - Space telescopes

# ALTAIR/NIRI

- **Near InfraRed Imager and Spectrograph**
- Single-conjugate adaptive optics (AO) system ALTAIR
- Gemini-North
- *JHK* broad band filters
- 22 arcsec x 22 arcsec
- 0.0219 arcsec/pixel
- FWHM ~ 0.1 arcsec
- Natural guide star (NGS) or laser guide star (LGS) modes
- Sample of 8 LIRGs
  - < 100 Mpc
  - SN rate ~ 1 yr<sup>-1</sup>



[www.noao.edu](http://www.noao.edu)

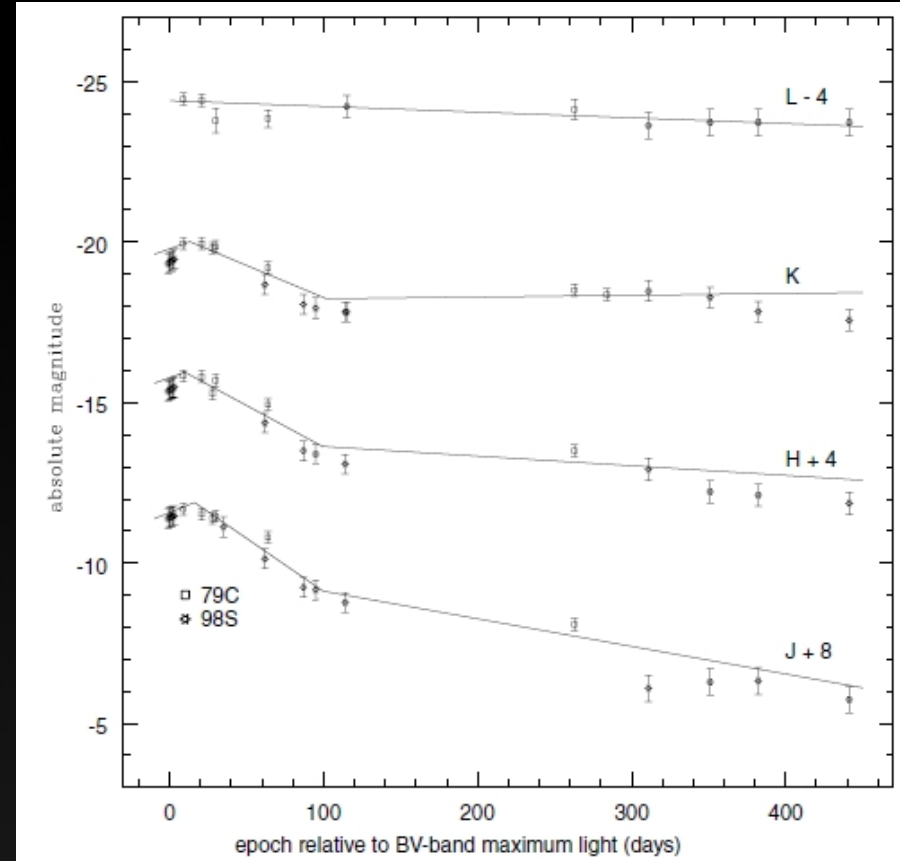
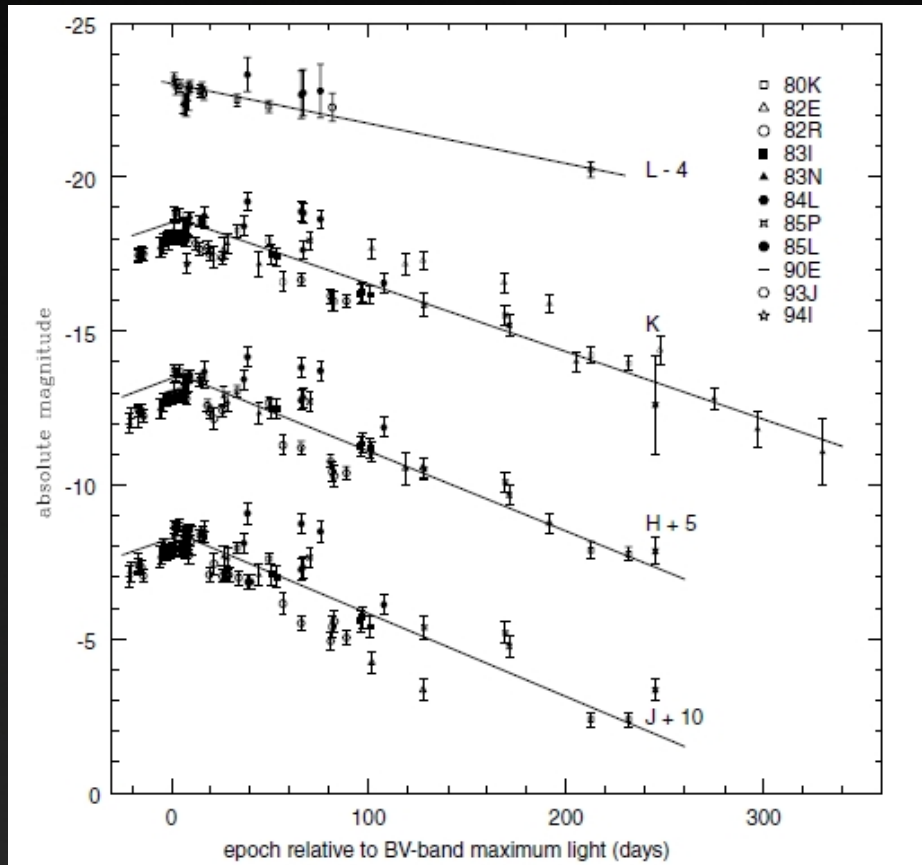


# Near-IR light curve templates

- **Mattila & Meikle (2001):**

- 'normal' & 'slowly declining SNe'

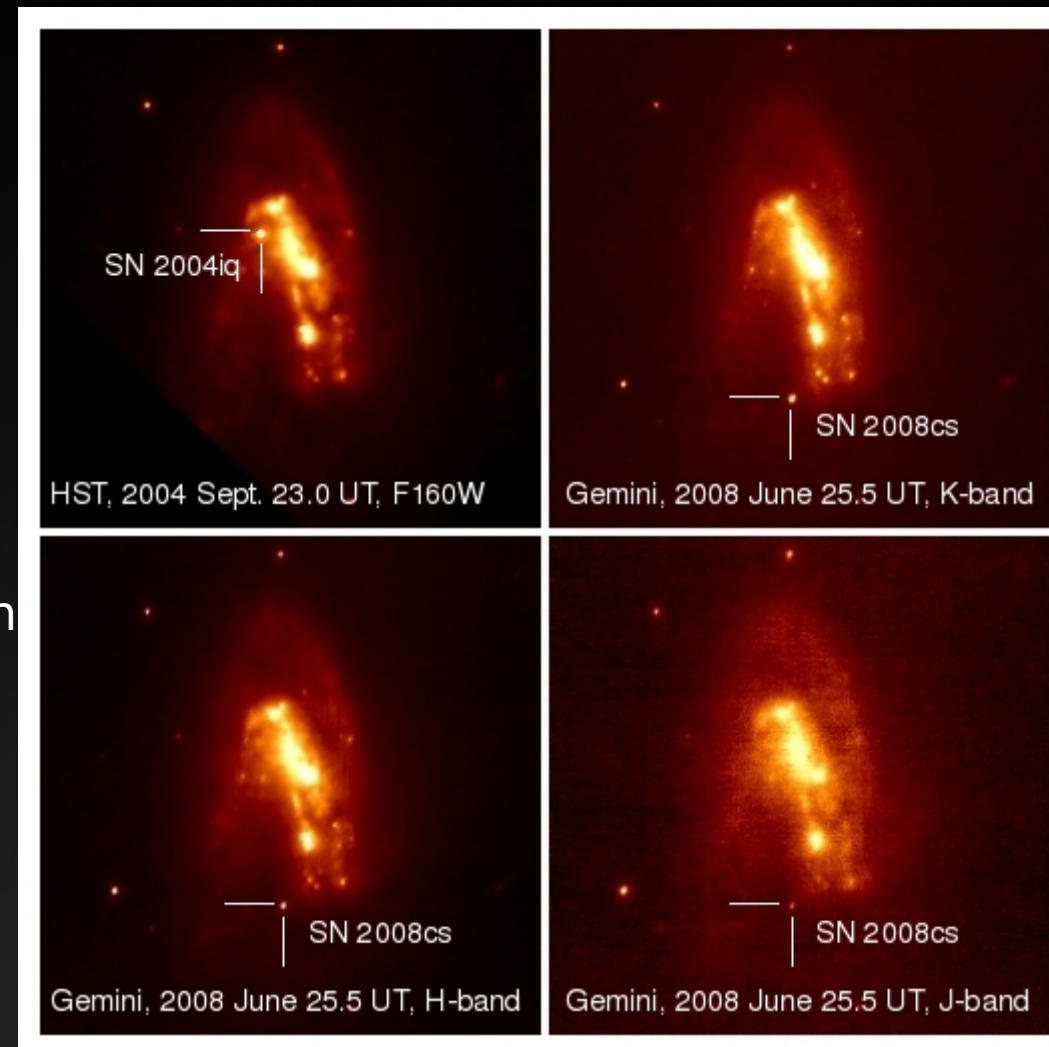
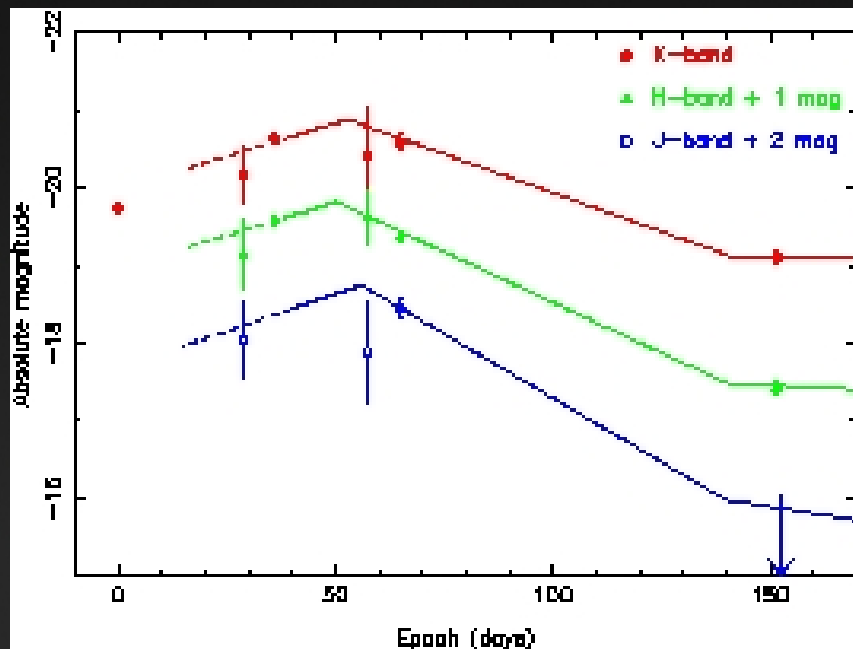
- 'normal' template based on several Type Ib/c SNe (also some Type II-P and Type II-L data)



- 'slowly declining' template based on two SNe: Type II<sub>n</sub> SN 1998S and II-L SN 1979C
- However, volume limited LOSS sample:
  - 53% CCSNe Type II-P SNe
- Well sampled Type II-P *JHK* light curve of SN 1999em (Krisciunas et al. 2009) as a temporary template

# SN 2004iq & SN 2008cs in IRAS 17138-1017

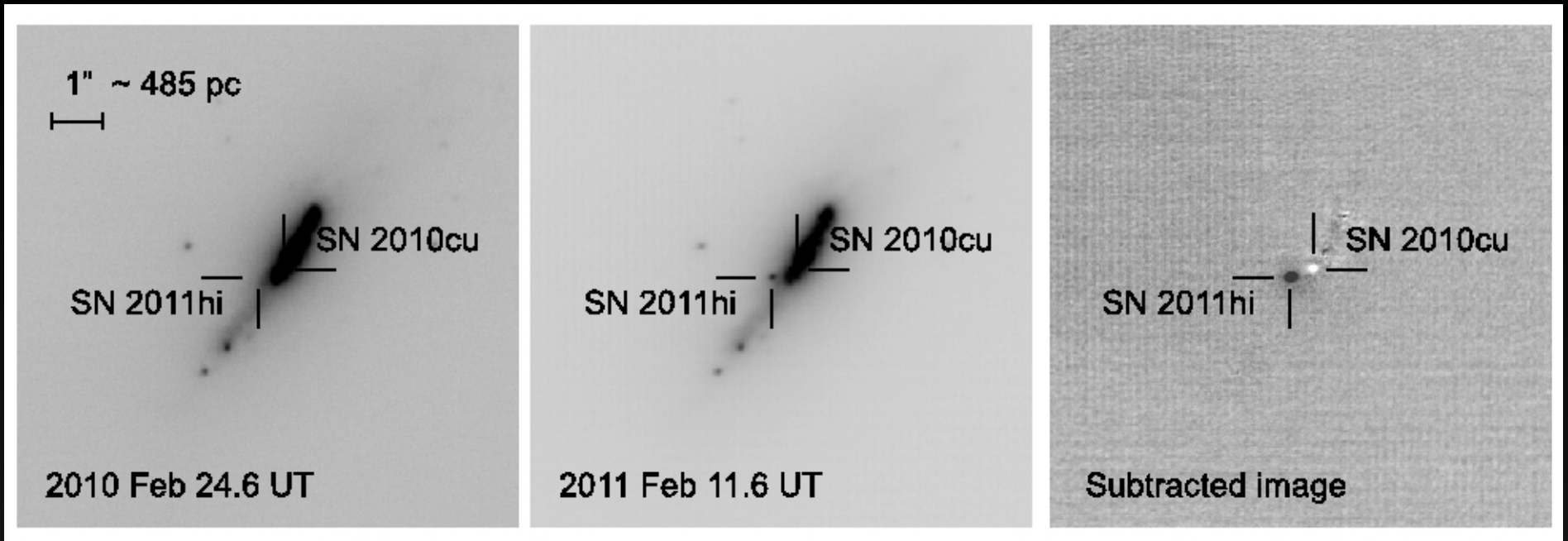
- Kankare et al. (2008), *ApJ*, 689, L97
- SN 2004iq: SN in HST archive images
  - Host galaxy extinction of  $A_V \sim 0-4$  mag?
- SN 2008cs: First LGS SN discovery
  - Total line-of-sight extinction of  $A_V \sim 18$  mag
  - Rieke & Lebofsky (1985) extinction law



- Expected SN rate of IRAS 17138-1017  $\sim 0.7$  SNe  $\text{yr}^{-1}$

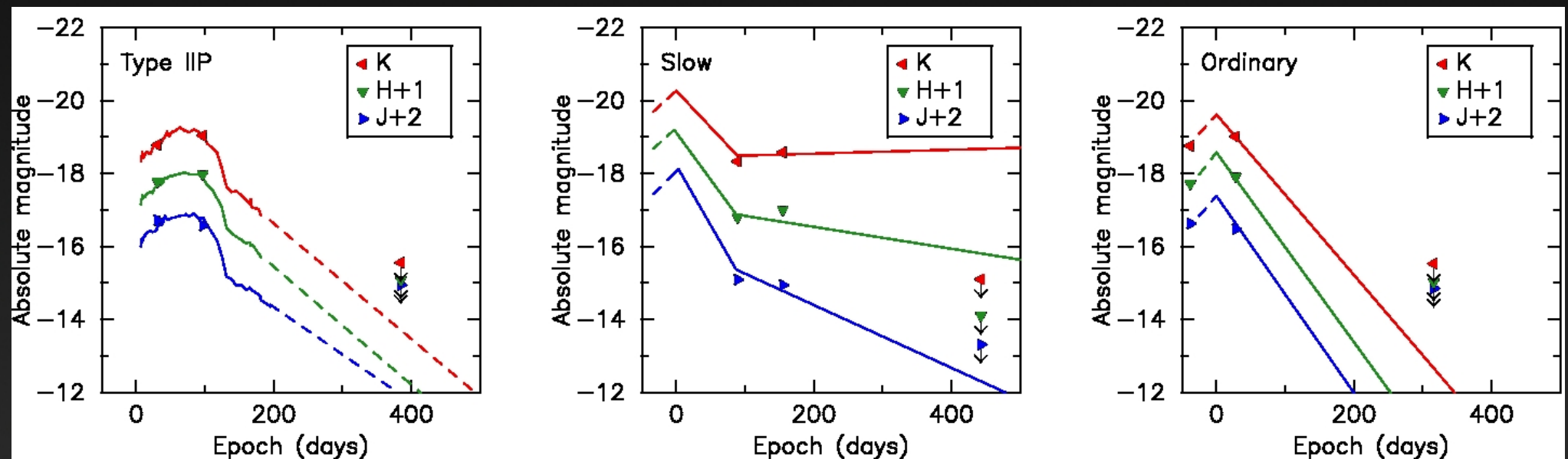
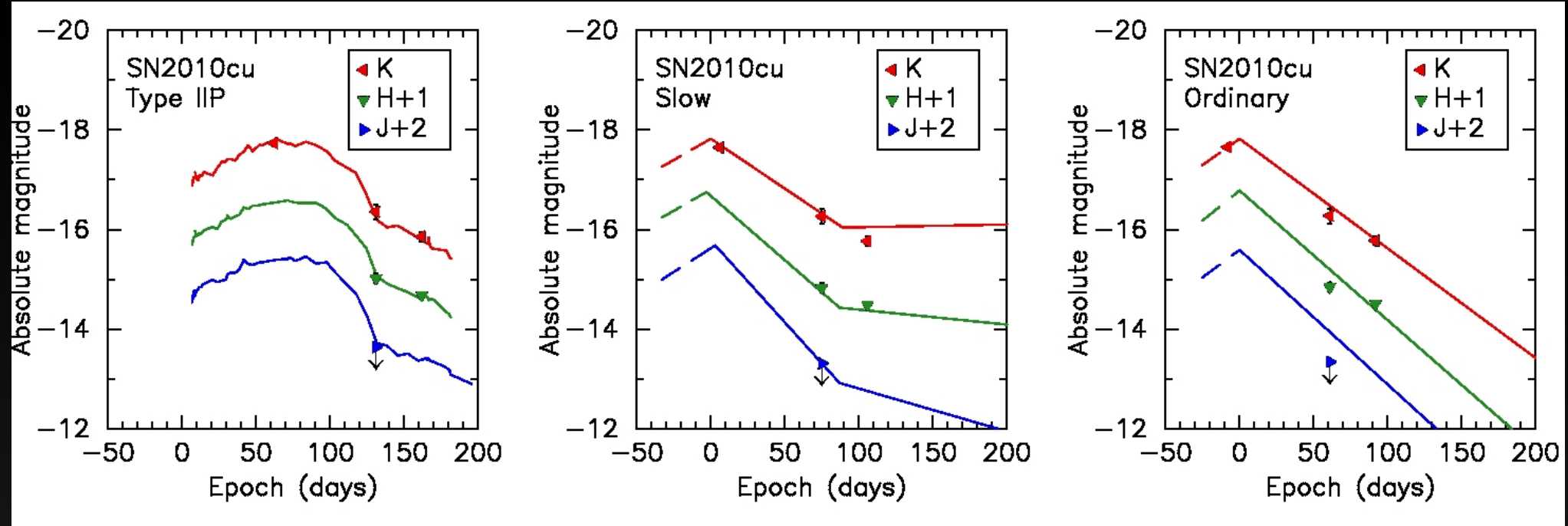


# SN 2010cu & SN 2011hi in IC 883



- [Kankare, Mattila, Ryder et al. \(2012\), ApJ, 744, L19](#)
- Radio paper on IC 883 with additional ALTAIR/NIRI data, [Romero-Cañizales et al. \(2012\), A&A in press](#)
- Expected SN rate of IC 883  $\sim 1.3 \text{ SNe yr}^{-1}$
- SN 2010cu at  $0.4''$  (180 pc) and SN 2011hi at  $0.8''$  (380 pc) from the nucleus (projected distance)

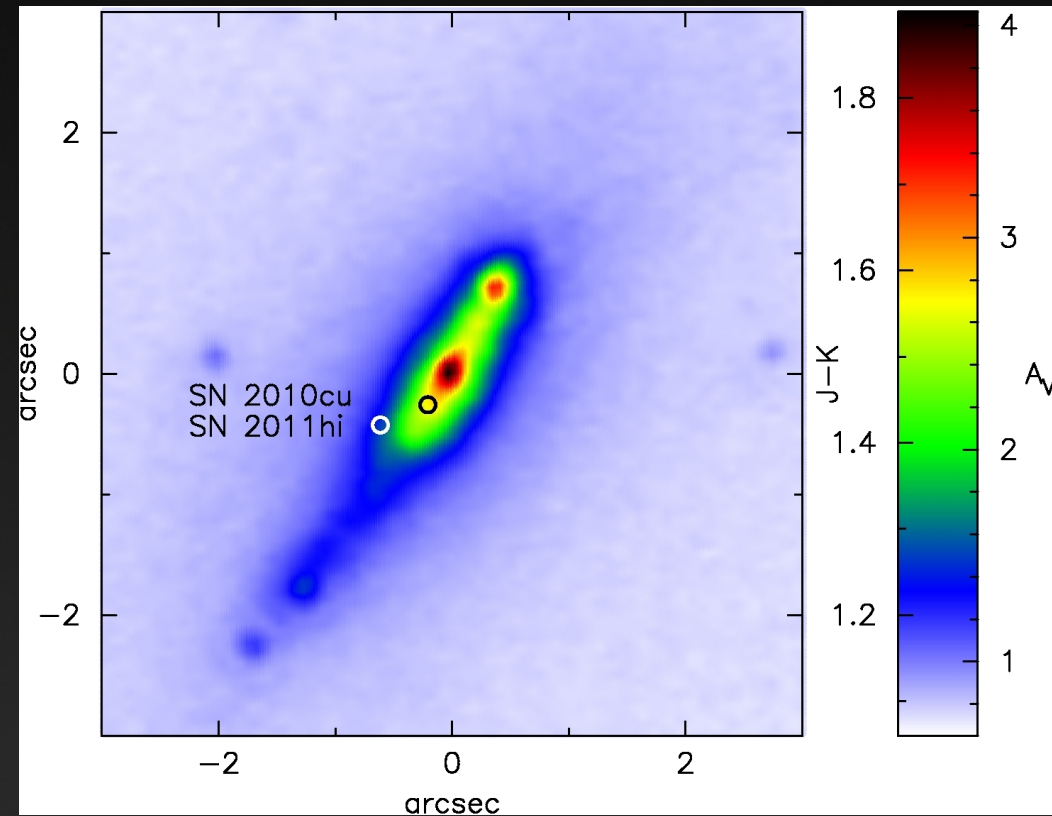
# SN 2010cu & SN 2011hi in IC 883



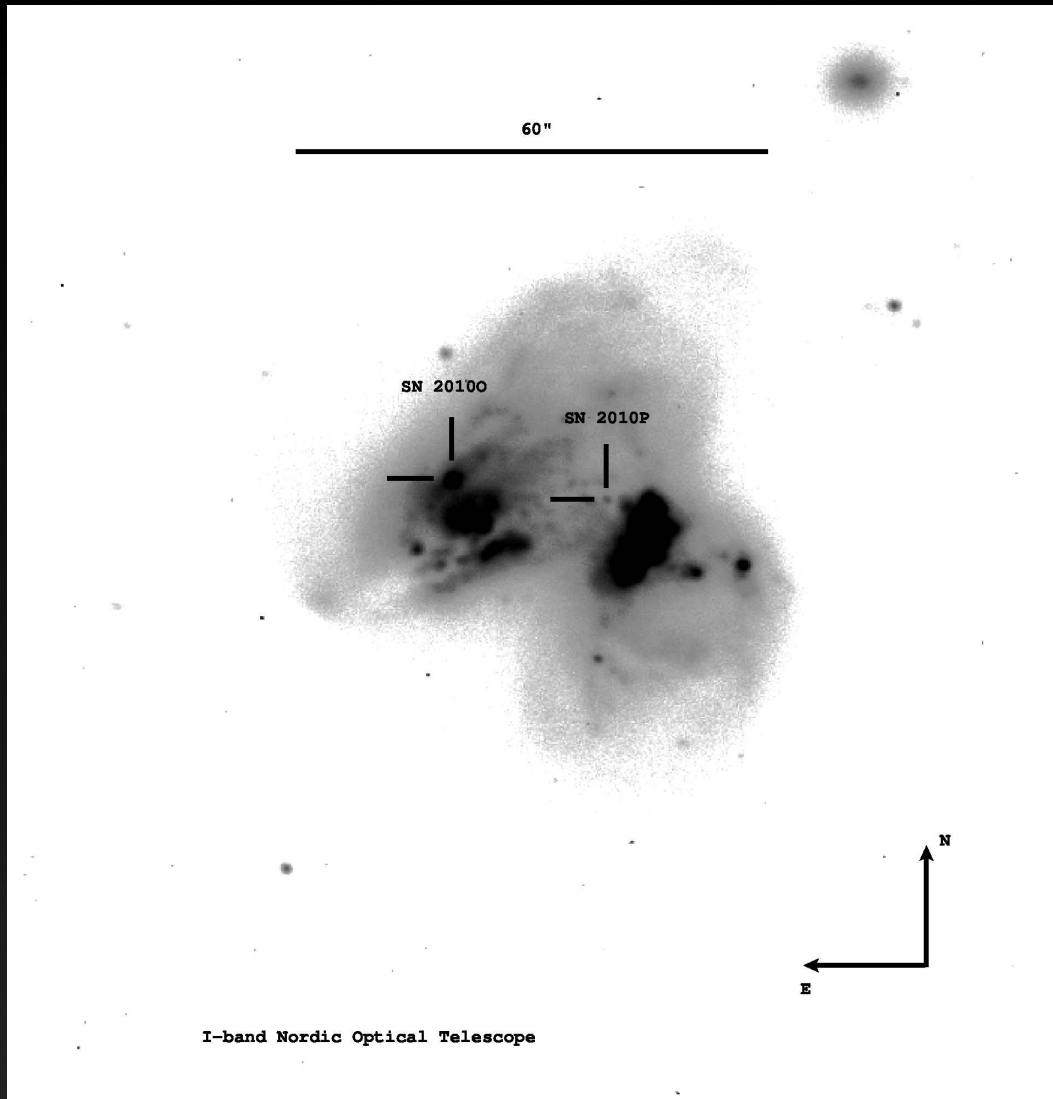
# SN 2010cu & SN 2011hi in IC 883

Template	$A_V$ (mag)	$t_0$ (days)	$C$ (mag)	$\tilde{\chi}^2$	$A_V$ (mag)	$t_0$ (days)	$C$ (mag)	$\tilde{\chi}^2$
SN 2010cu	Cardelli law				Calzetti law			
ordinary	0.1	-8	+0.80	11	0.1	-8	+0.80	11
slow	0.0	6	+2.20	8.5	0.0	6	+2.20	8.5
IIP	1.3	62	+0.10	1.7	1.0	62	+0.15	1.7
SN 2011hi								
ordinary	6.8	-38	-1.35	5.0	4.7	-38	-1.00	4.4
slow	0.0	89	-0.25	27	0.0	89	-0.25	27
IIP	7.0	31	-1.65	9.4	5.0	31	-1.30	8.8

- $J$ - $K$  HST/NICMOS archive data colour map
  - Upper limit for the line-of-sight extinction
  - Localized higher extinction possible
- SN 2010cu most consistent with 'slowly declining' and IIP templates
- New data on 31st of January 2012. No clear detection of SN 2011hi → likely a Type IIP



# SN 2010O & SN 2010P in Arp 299

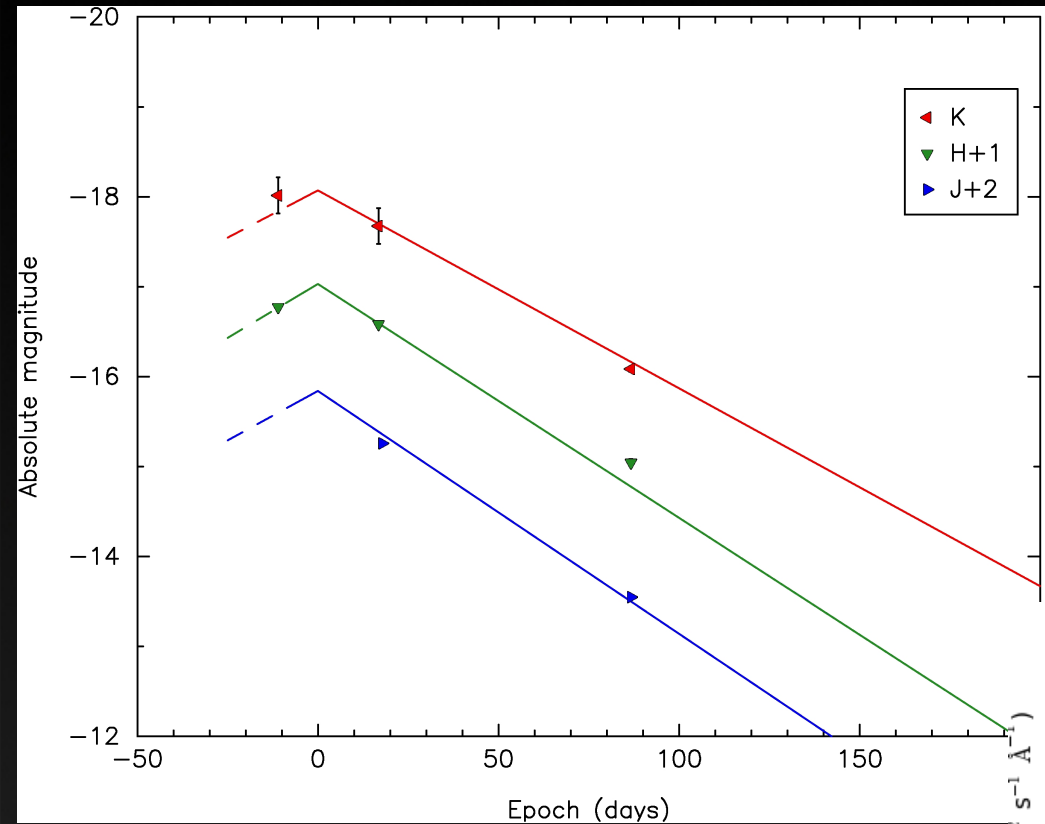


- SN 2010O discovered by [Newton et al. \(2010\)](#), CBET 2144
- SN 2010P discovered by [Mattila & Kankare \(2010\)](#), CBET 2145
- SN 2010O typed as Ib by [Mattila, Kankare, Datson & Pastorello \(2010\)](#), CBET 2149
- SN 2010P typed as Ib/IIb by [Ryder, Mattila, Kankare & Perez-Torres \(2010\)](#), CBET 2189

- Expected SN rate of Arp 299  $\sim 2 \text{ SNe yr}^{-1}$

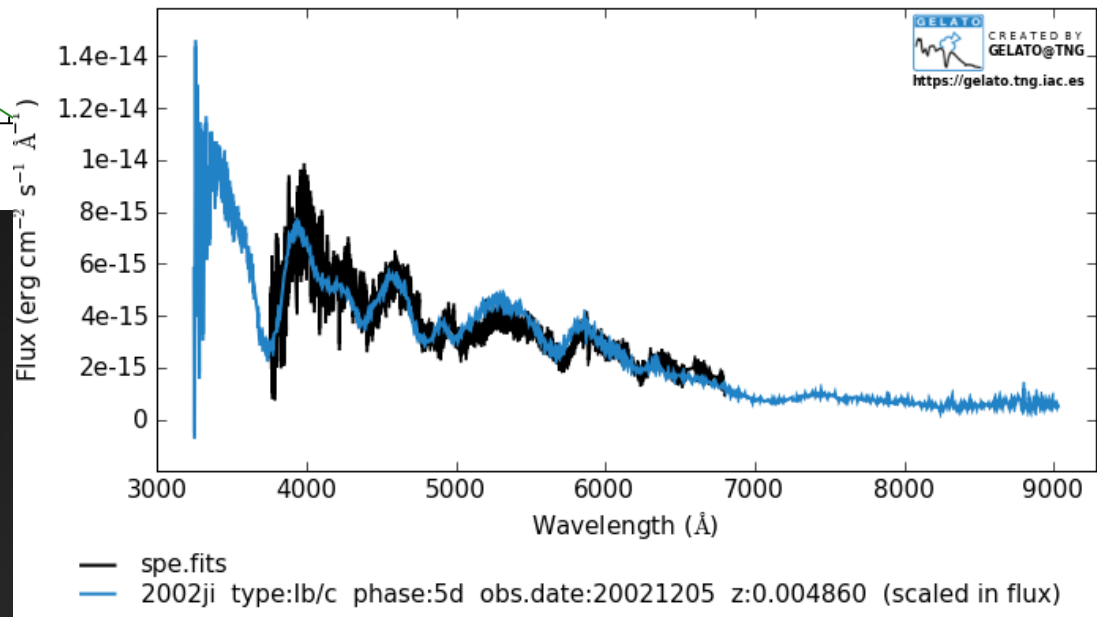


# SN 2010O in Arp 299

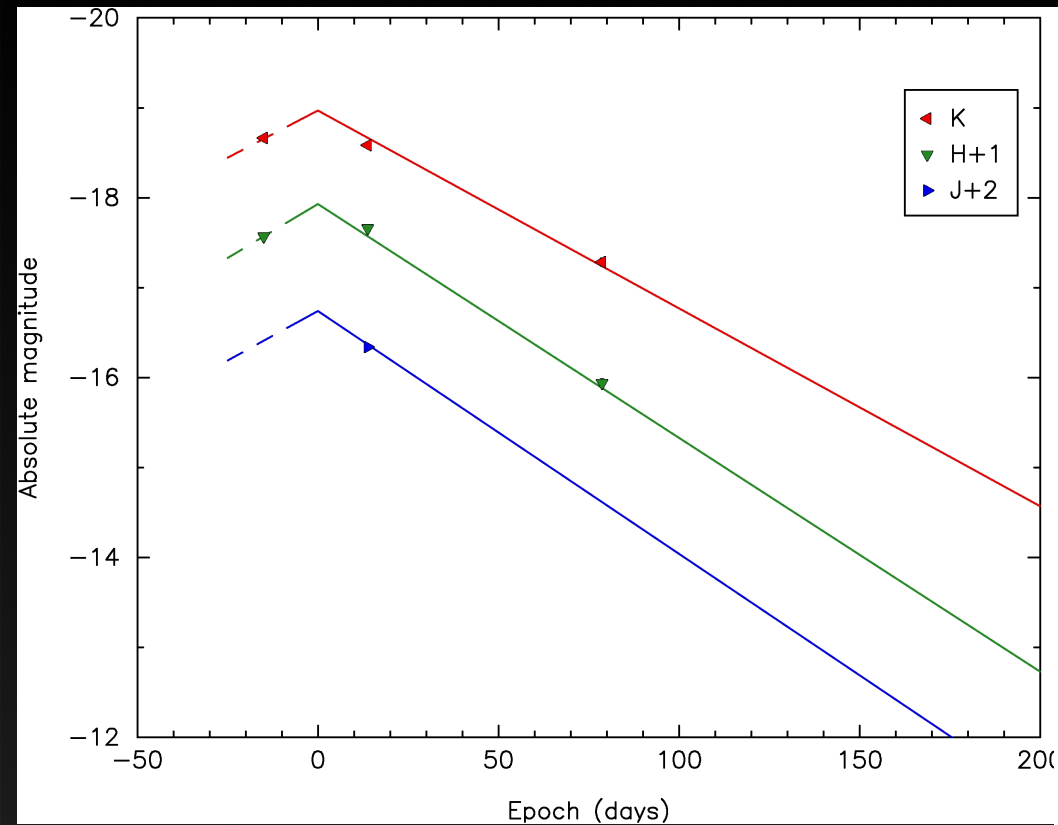


- 'normally' declining template
- $A_V = 2.7$  mag
- Discovered on day -11 respective to the peak
- Scaling constant 0.55 mag

- Spectroscopically typed as a Ib SN
- Consistent with the spectroscopic typing and extinction estimate of 2 mag derived based on the spectrum

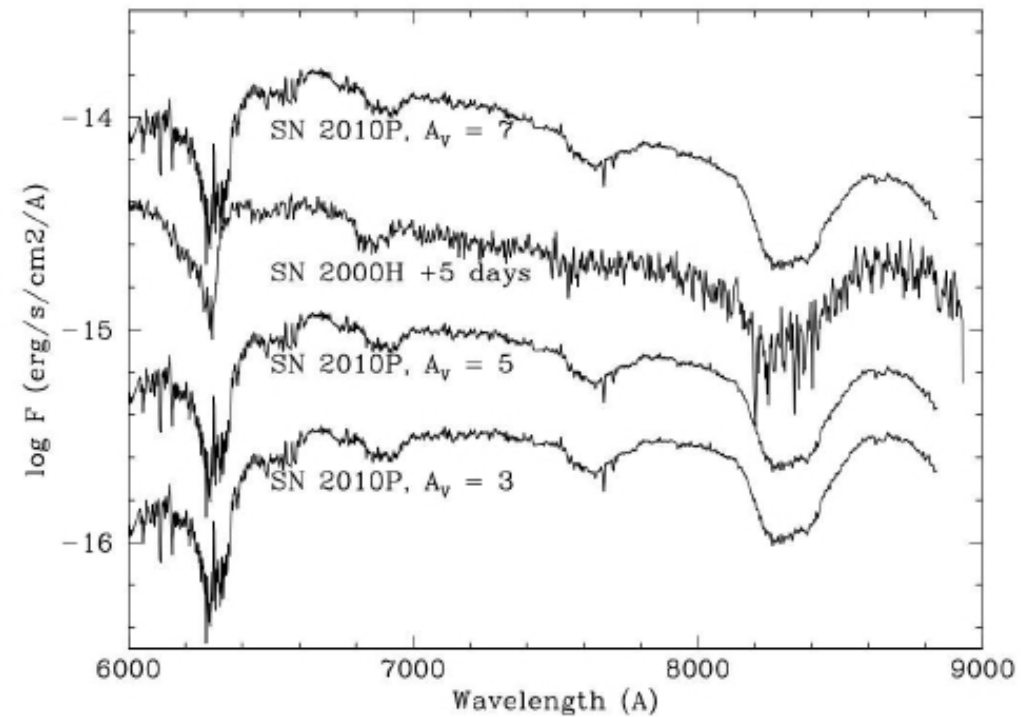


# SN 2010P in Arp 299



- 'normally' declining template
- $A_V = 7.0$  mag
- Discovered on day -15 relative to the peak
- Scaling constant -0.35 mag

- Spectroscopically typed as a Ib/Ib SN
- Consistent with the spectroscopic typing and extinction estimate of 5-7 mag derived based on the spectrum



Mattila, Dahlen, Efstathiou, Kankare et al.  
(submitted to ApJ)



# VLT/NaCo

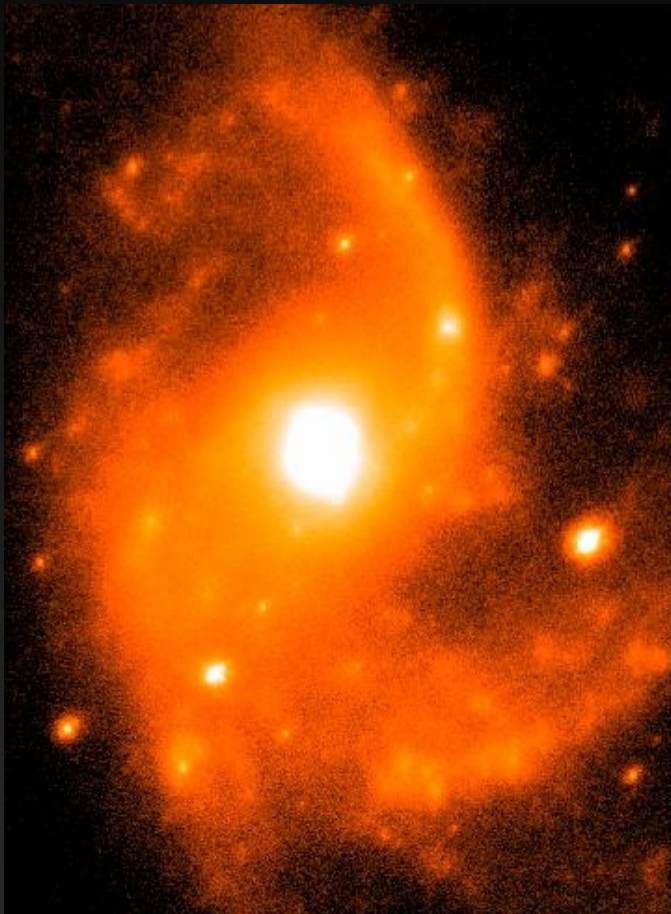
- LIRG SN project continuing with the VLT/NaCo in the period 89
- Nasmyth Adaptive Optics System (NAOS) Near-Infrared Imager and Spectrograph (CONICA) in UT4
- Single-conjugate AO system with a natural guide star



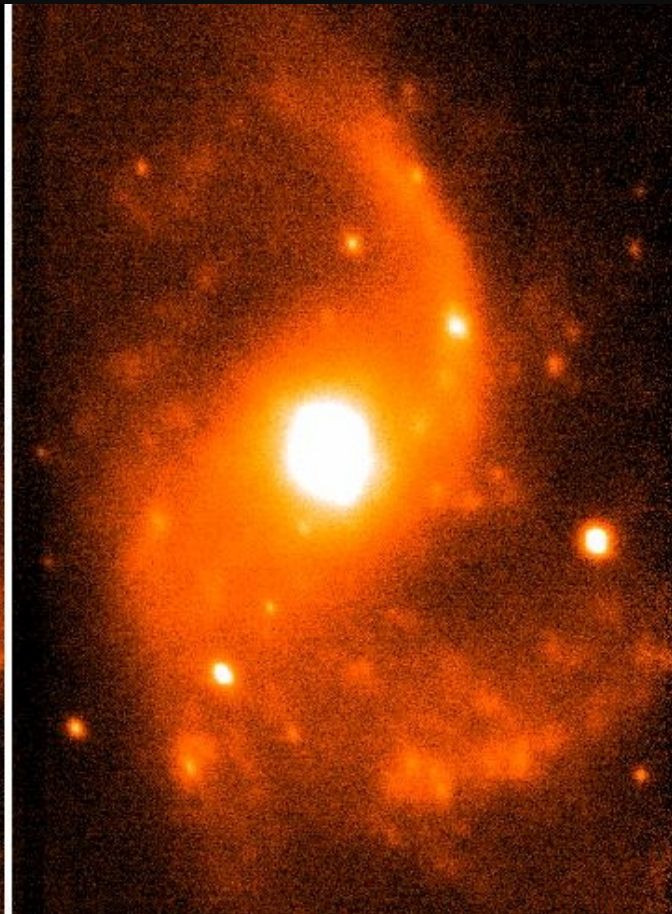
[www.eso.org](http://www.eso.org)

# VLT/NACO

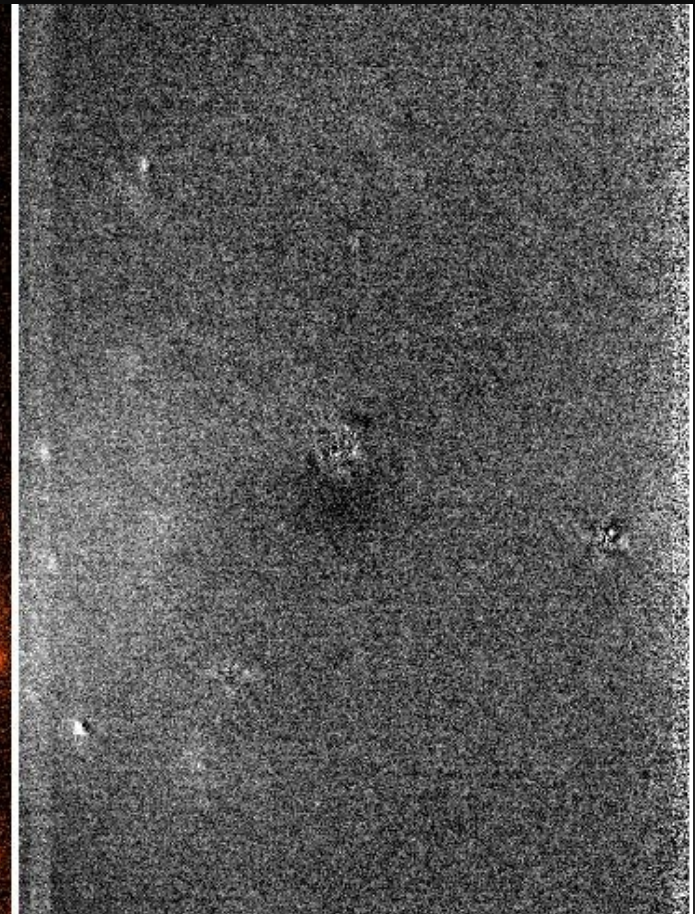
- Example of VLT/NACO data
  - ESO 221-IG 010
  - Expected SN rate of  $\sim 0.4$  SNe yr<sup>-1</sup>
  - Smooth subtraction, no SNe



2011 February 14<sup>th</sup>



2012 May 21<sup>st</sup>



subtraction



# Summary

- Increased statistics important for mapping SN populations and extinctions distributions of LIRGs
- 6 SNe discovered in 8 LIRGs within 4 years
- Line-of-sight extinctions and types estimated based on near-IR light curves, colours and magnitudes

SN	Type	$A_V$
2004iq	?	0 - 4 mag
2008cs	IIIn/II-L	14 - 16 mag
2010O	Ib	2 - 3 mag
2010P	Ib/IIb	5 - 7 mag
2010cu	II	0 - 1 mag
2011hi	IIP	5 - 7 mag

- Next generation multi-conjugate adaptive optics systems and large sample size will dramatically increase the number of discovered CCSNe in LIRGs

Thank you for your attention

