

Hot accretion flow in X-ray binaries: spectral and timing evidence



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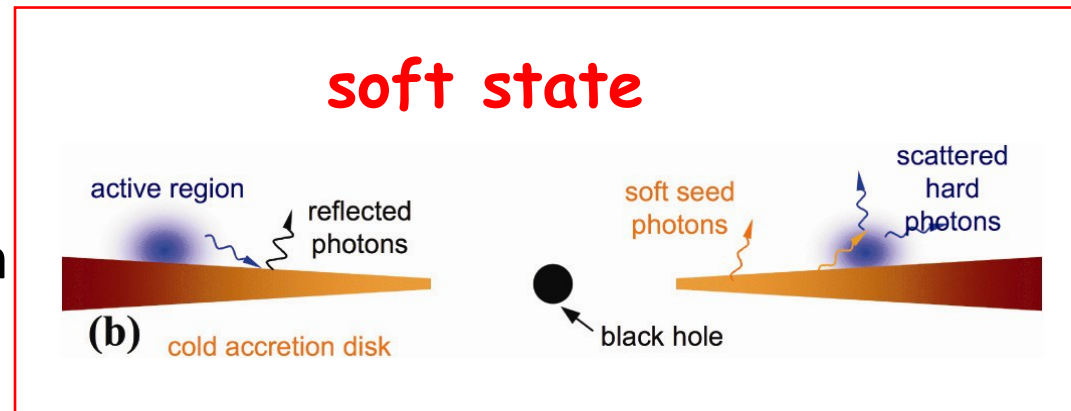
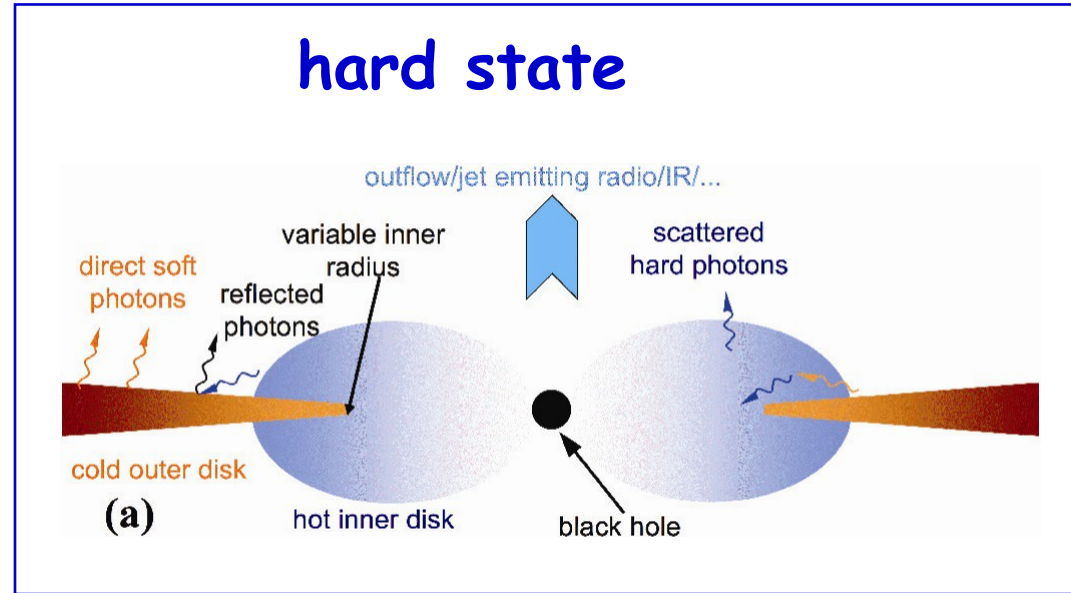
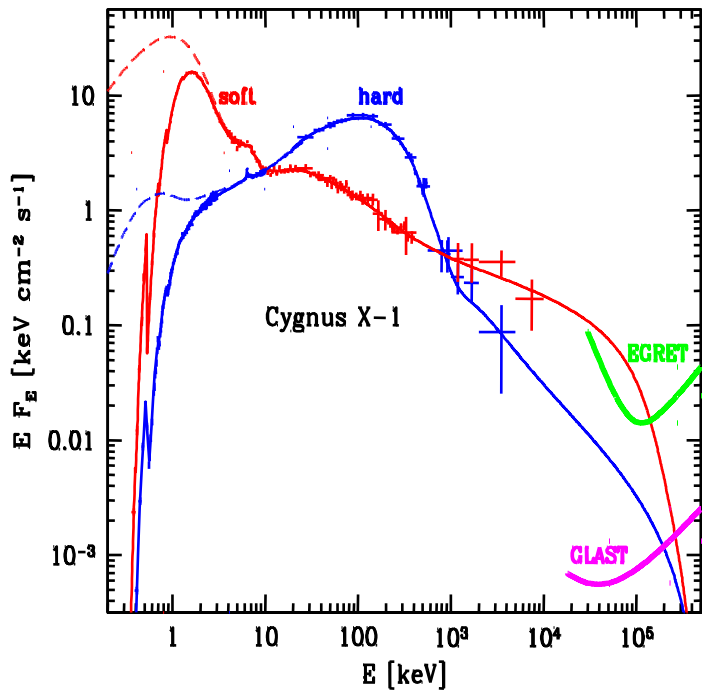
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Outline

- Black hole spectra: radio to X-rays
- Optical/X-ray variability
- Spectral and timing modelling

OBSERVATIONS

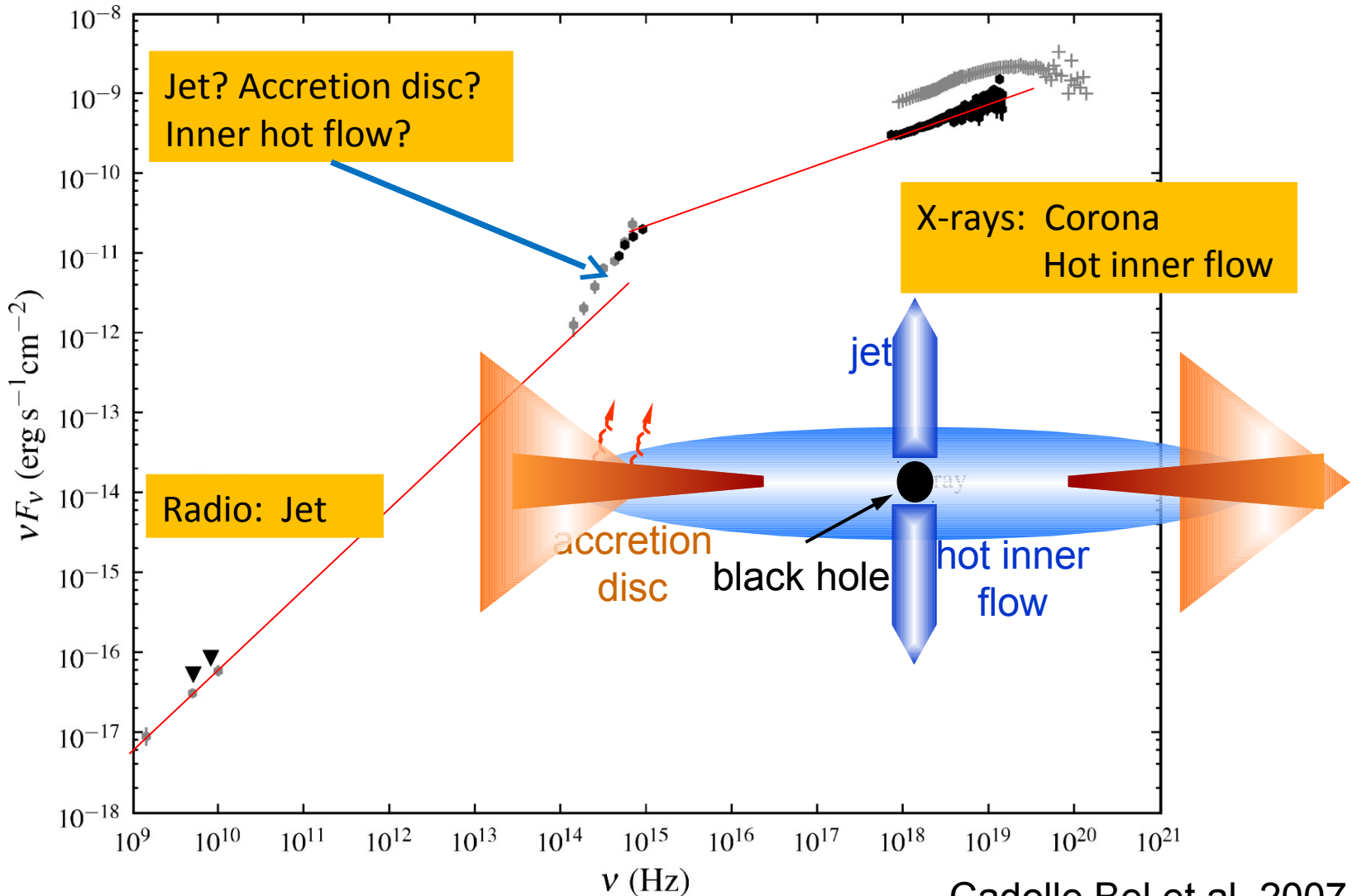
X-ray spectra and geometry



Hard state - standard cold outer disk + hot inner flow?

Soft state - standard accretion α -disk, plus corona?

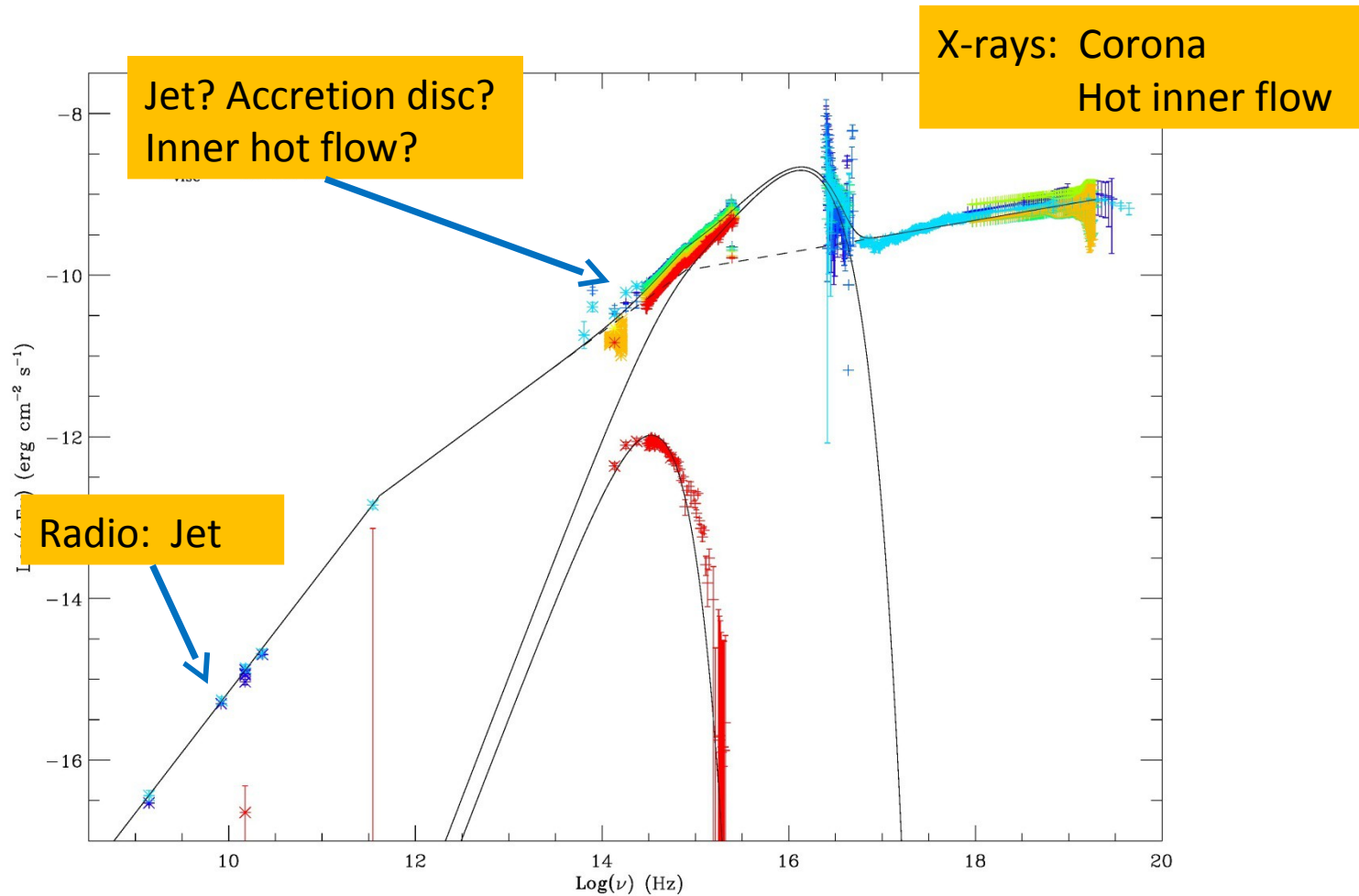
Broadband spectra of LMXBs



Swift J1753.5-0127

Cadolle Bel et al. 2007,
Durant et al. 2009

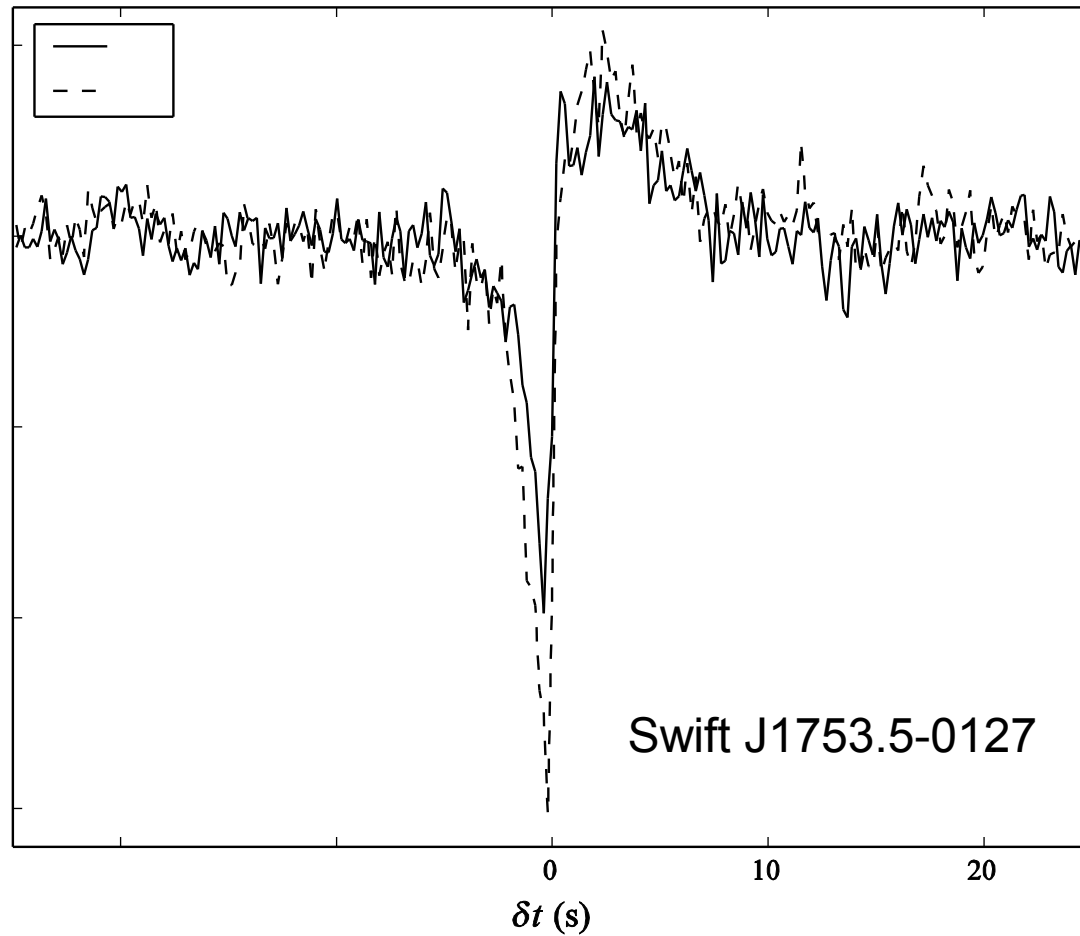
Broadband spectra of LMXBs



XTE J1118+480

Chaty et al. 2003

Optical/X-ray cross-correlation



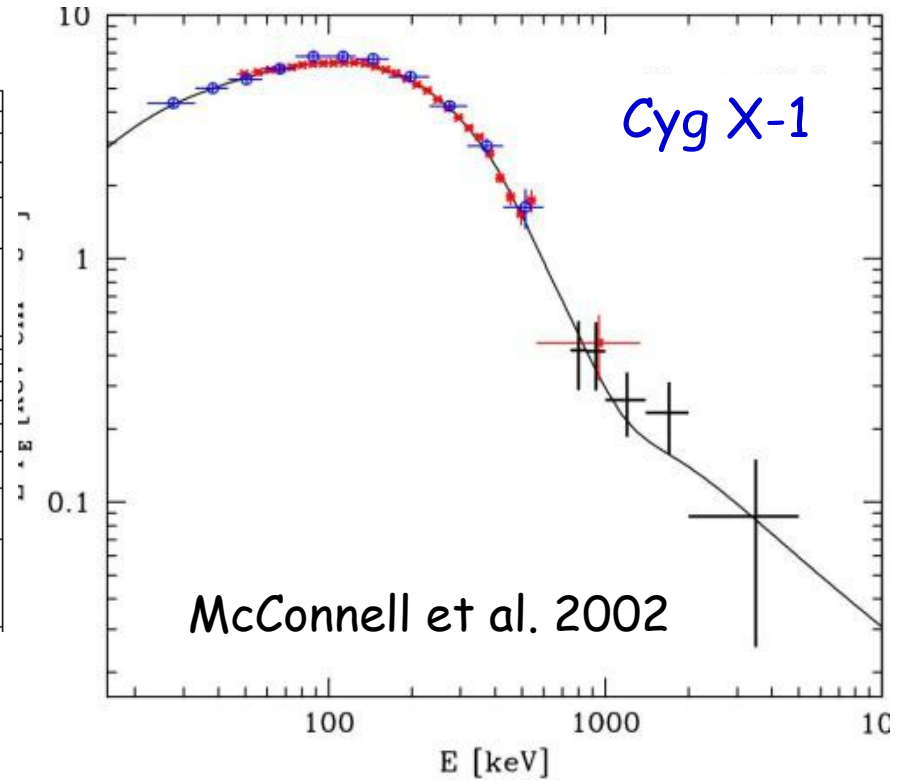
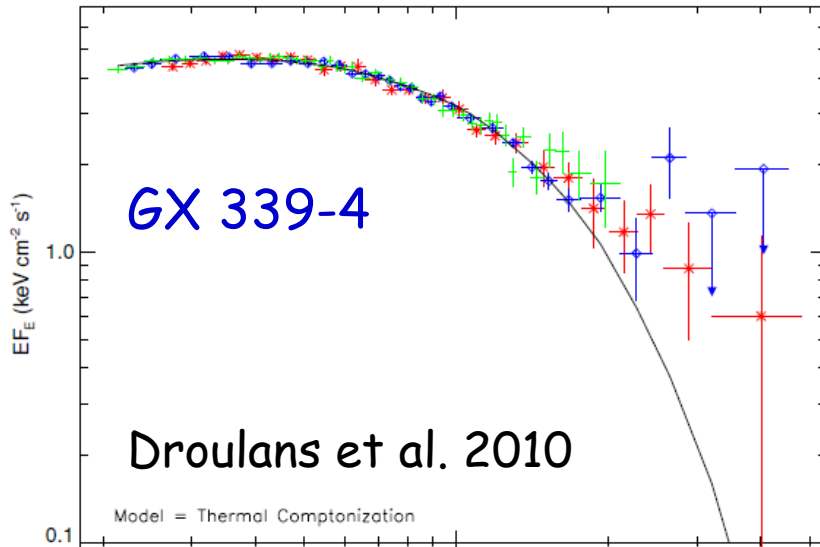
Durant et al. 2010

Observational aspects

- Spectrum: optical/infrared spectrum is inconsistent with being produced by standard accretion disc or the jet
- Timing: mysterious shape of the cross-correlation function

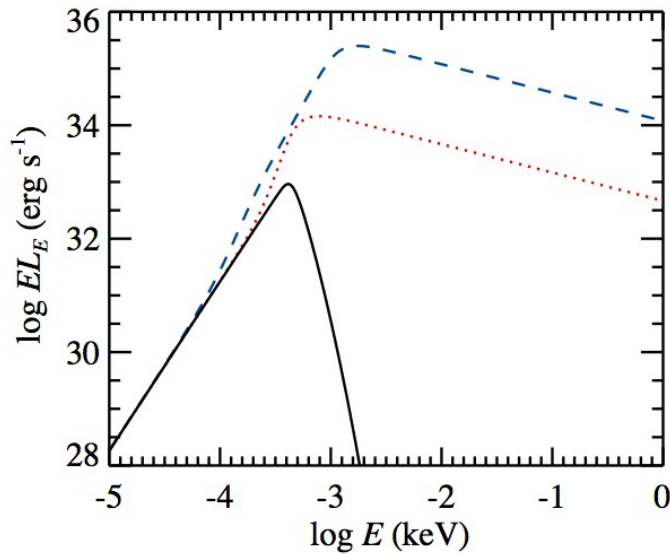
MODELING

Thermal Comptonization in the hard state



A weak non-thermal tail is present

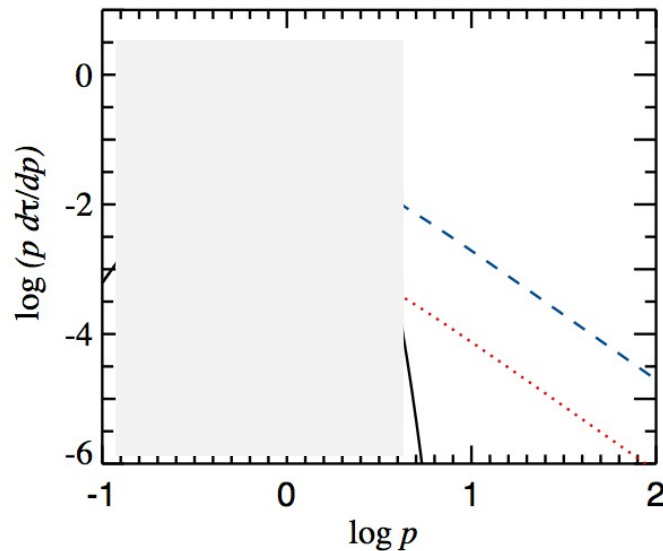
Synchrotron in hybrid plasma



Hybrid electrons, 1% energy
in the non-thermal component

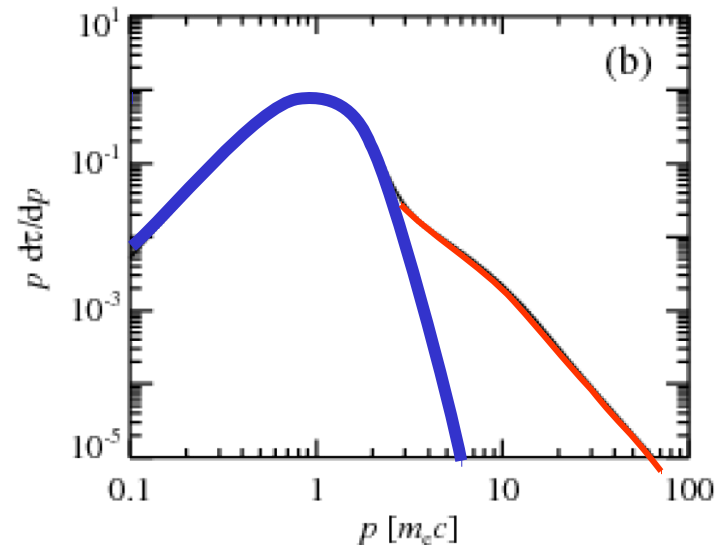
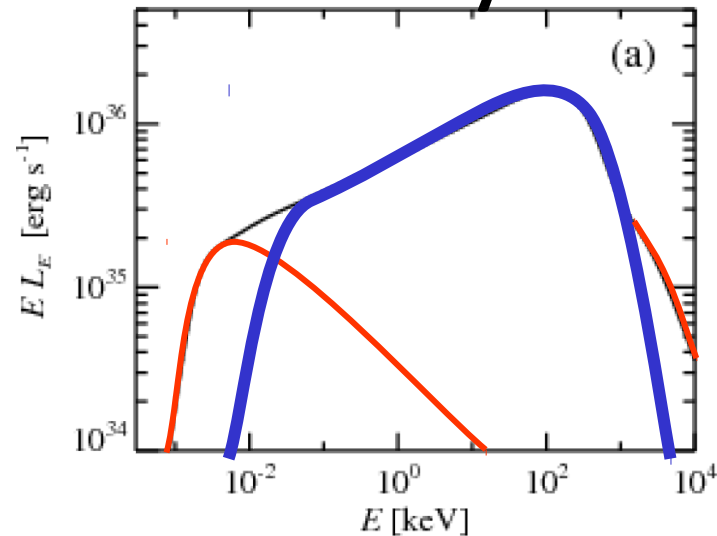
Hybrid electrons, 0.01% energy
in the non-thermal component

Thermal electrons, 100 keV



Synchrotron can be the main source of
seed photons for Comptonization

Synchrotron Self-Compton (SSC) mechanism in hybrid plasma

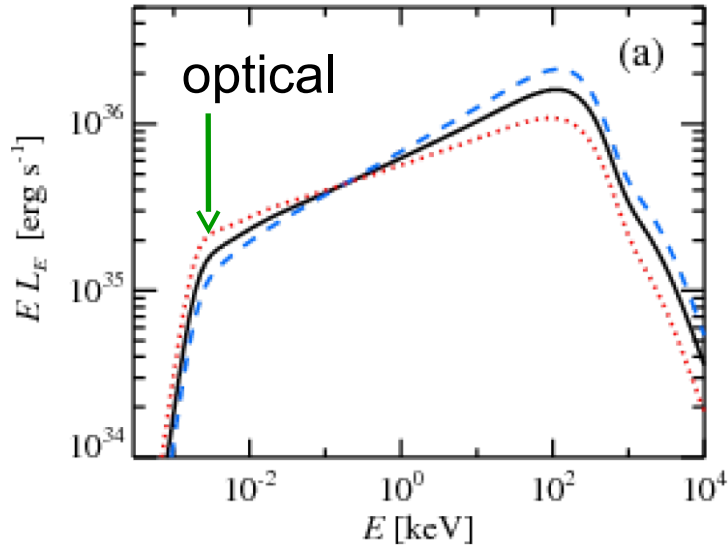


$$R = 9 \times 10^7 \text{ cm}$$

$$\tau = 1.0$$

$$B = 3 \times 10^5 \text{ G}$$

Synchrotron Self-Compton mechanism in hybrid plasma

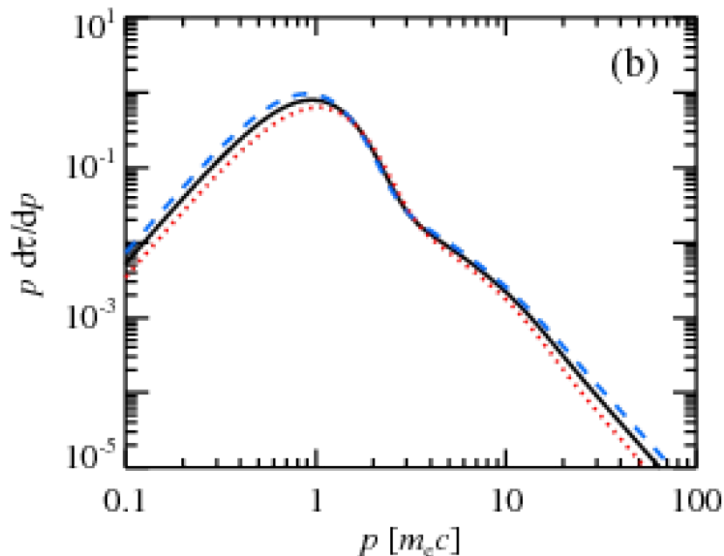


$$R \propto \dot{m}^{-4/3} \quad (\text{Rozanska \& Czerny 2000})$$

$$L \propto \dot{m}$$

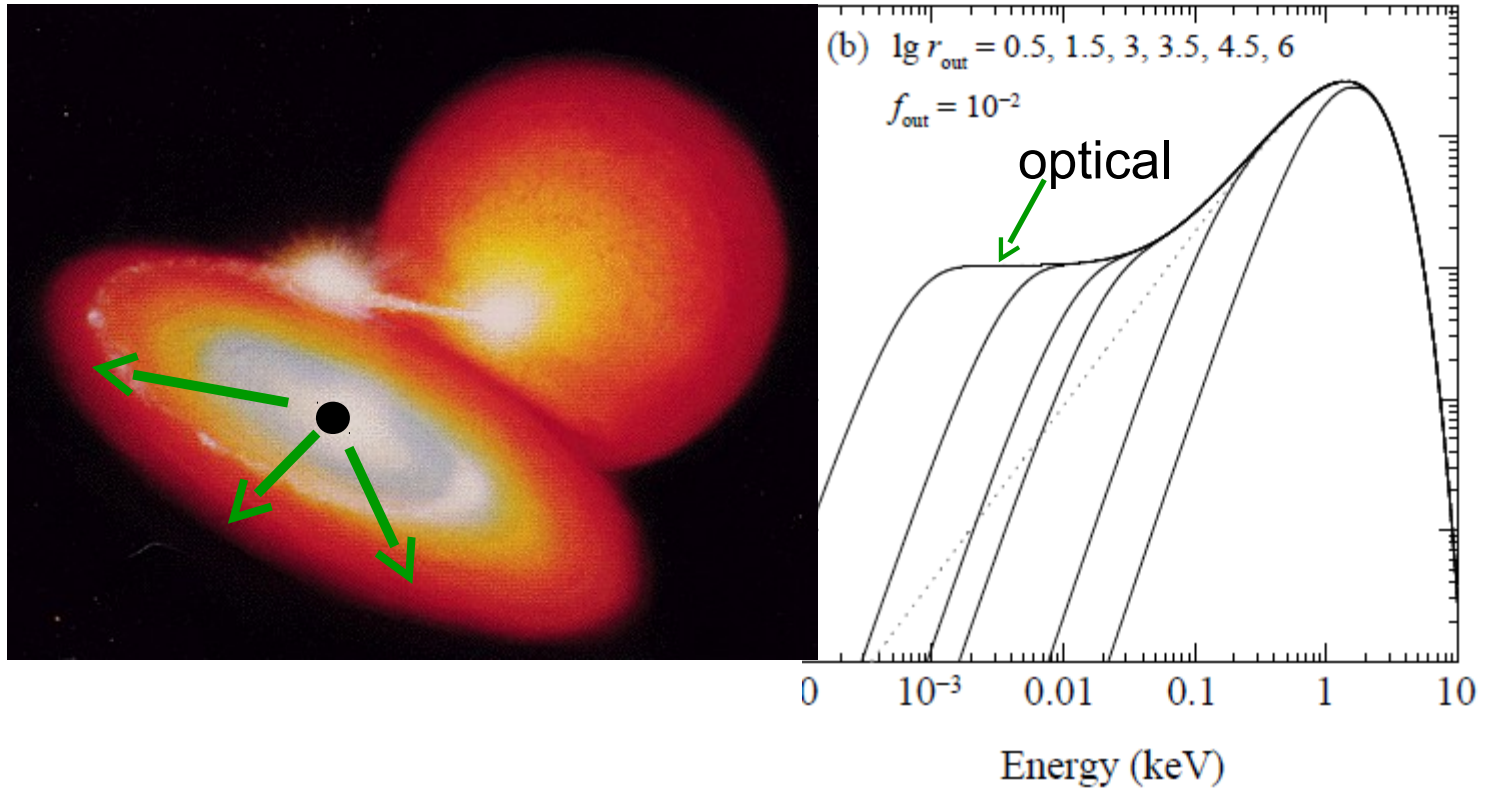
$$\tau \propto \dot{m}$$

$$B = \text{const}$$

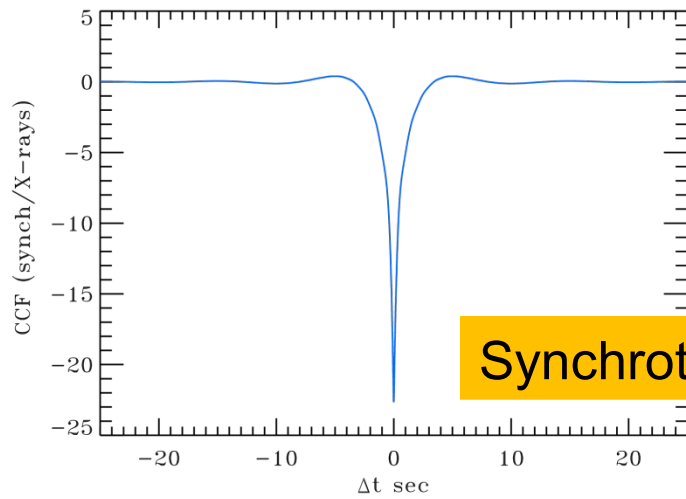


The optical and the X-rays are anticorrelated

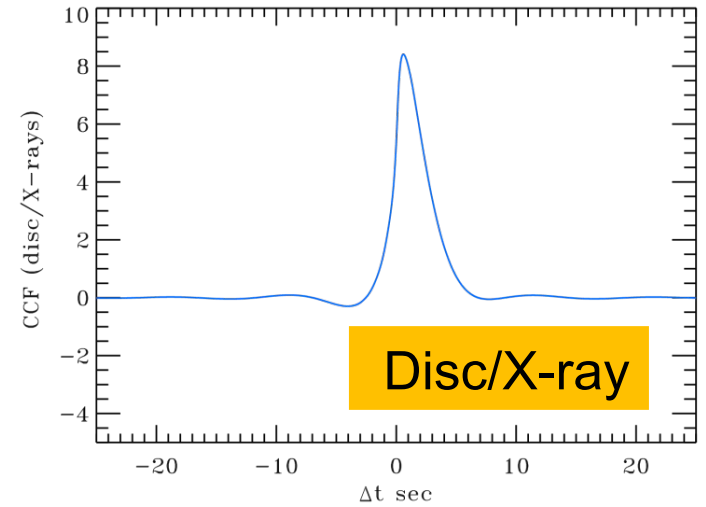
Irradiated discs



Optical/X-ray cross-correlation



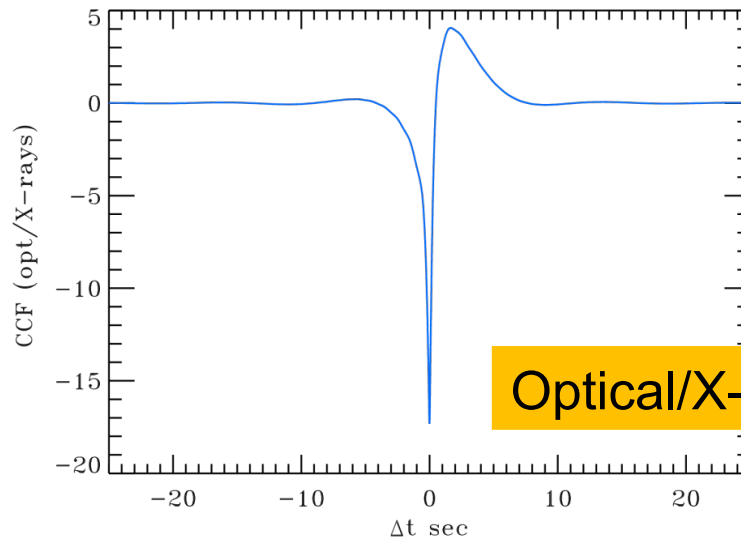
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Synchrotron/X-ray

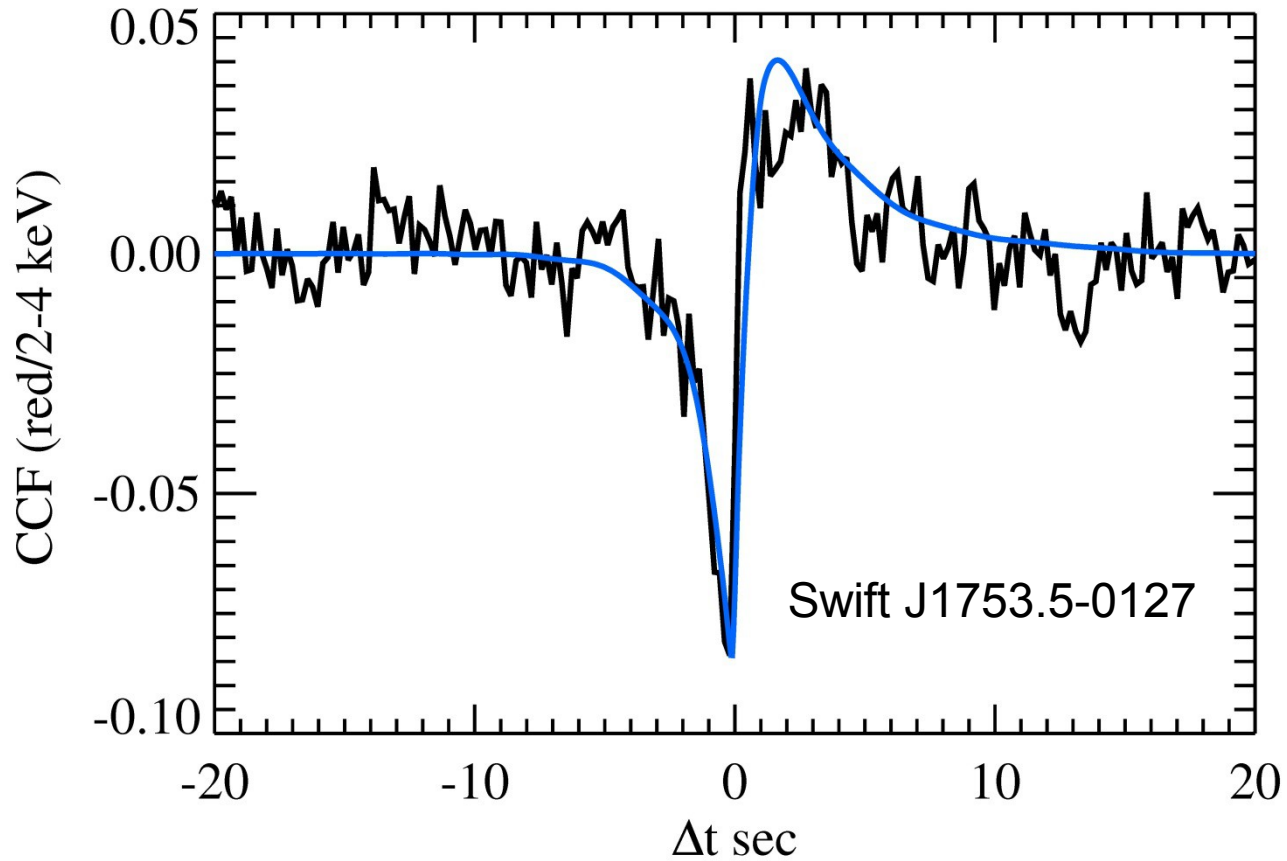
Disc/X-ray

=



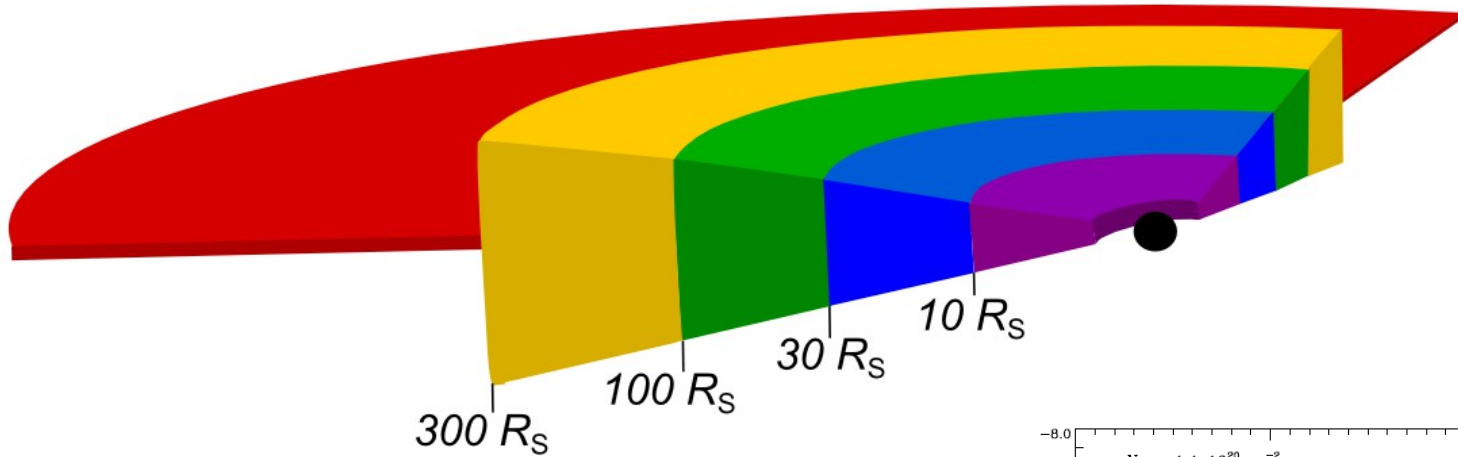
Optical/X-ray

Comparison with the data



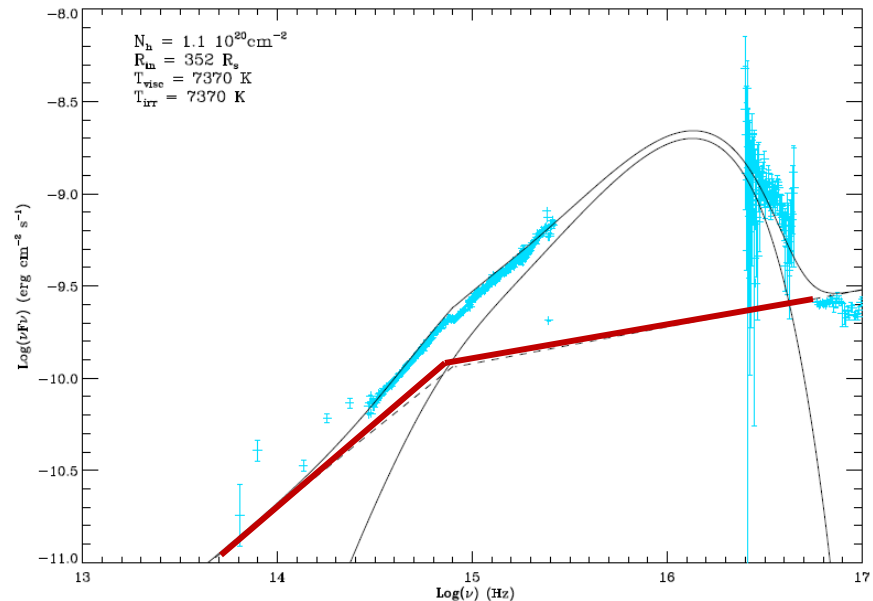
Data from Durant et al. 2010

Multi-zone hot accretion flow

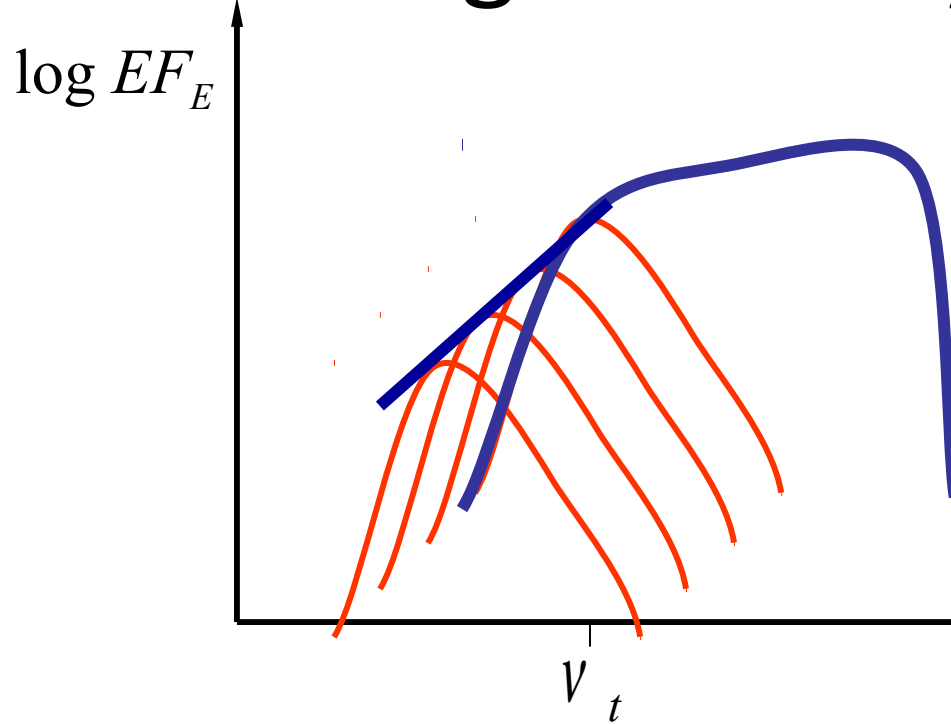


Hot inner flow can be divided into a number of zones. More compact zones have higher self-absorption frequency. The resulting OIR spectrum is flat

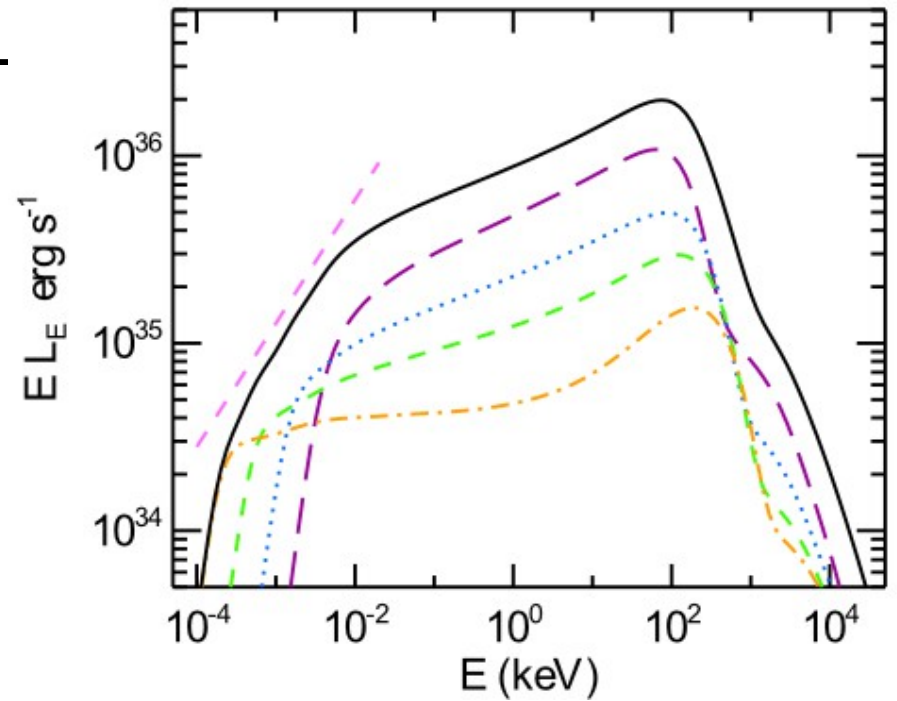
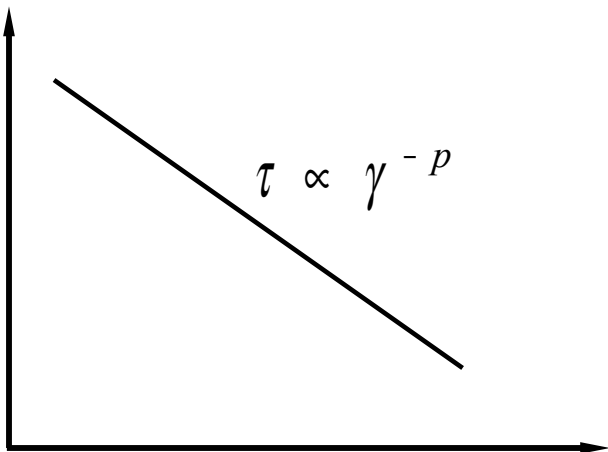
$$F_\nu \propto \nu^0$$



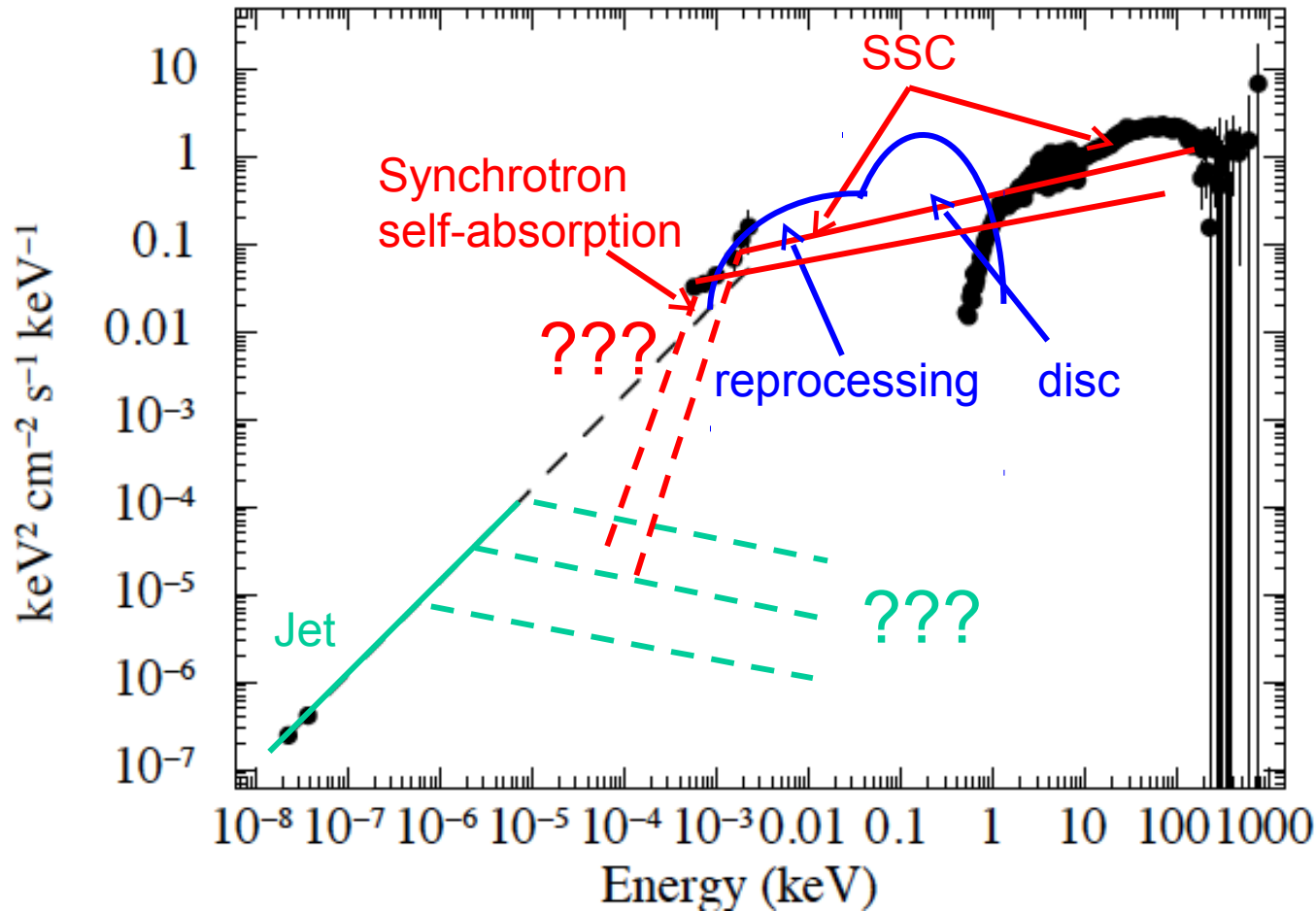
Inhomogeneous synchrotron source



$$\nu_t \propto B^{(p+2)/(p+4)} \tau^{2/(p+4)}$$



Broadband spectrum and multi-zone hot accretion flow



GX 339-4

Cadolle Bel et al. 2011

Conclusions

- The mysterious shape of the optical/X-ray CCF is explained by joint contribution of the [synchrotron + irradiated disc emission](#)
- Flat optical/infrared spectra can be explained in terms of [inhomogeneous SSC model](#)
- Question regarding the importance of the jet at these wavelengths remains to be studied